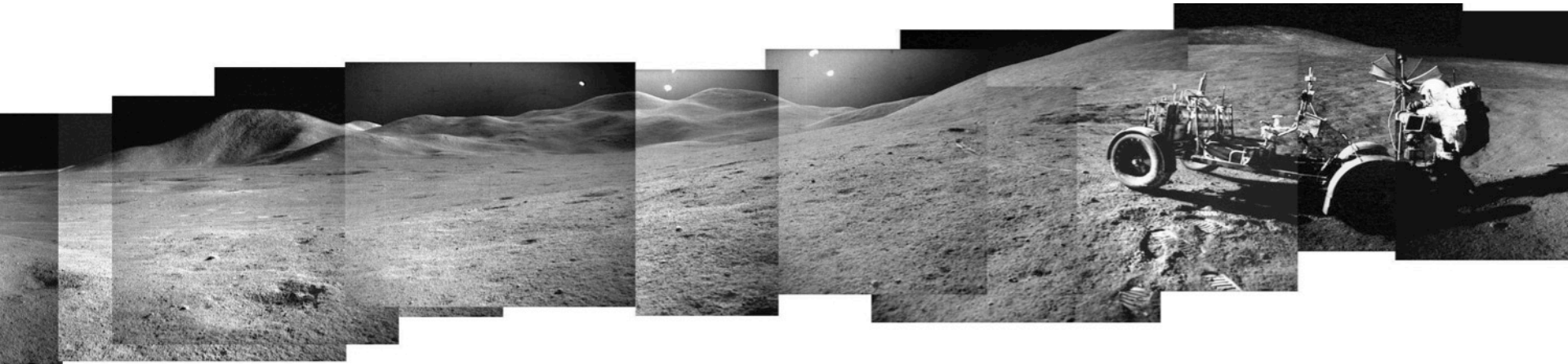


Elucidating the History of the Moon's Surface

High Spatial Resolution $^{40}\text{Ar}/^{39}\text{Ar}$ Geochronology of
Multigenerational Lunar Impact Melt Rocks

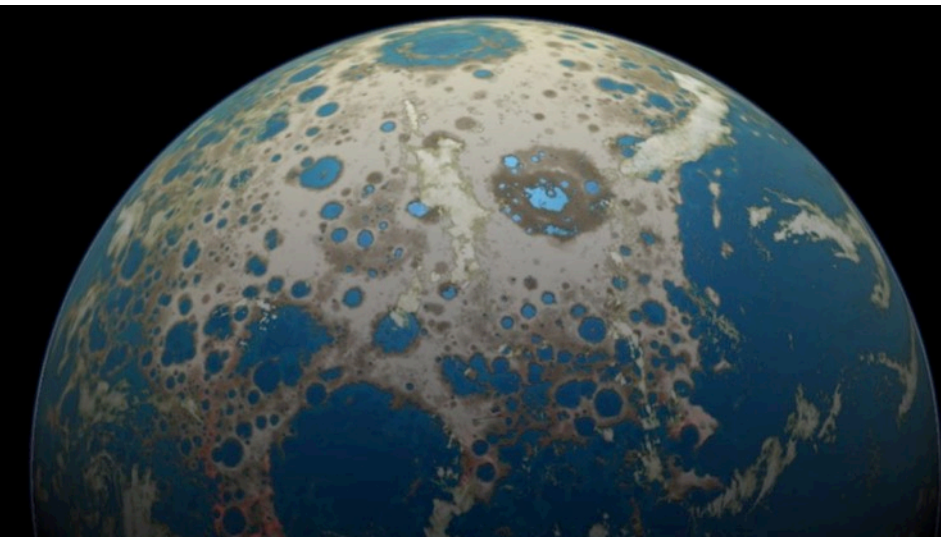


ASU School of Earth and
Space Exploration
Arizona State University

Hayden Miller
SESE Colloquium
March 17th, 2021

Why study the moon?

- Moon is an unique archive of early Solar System history



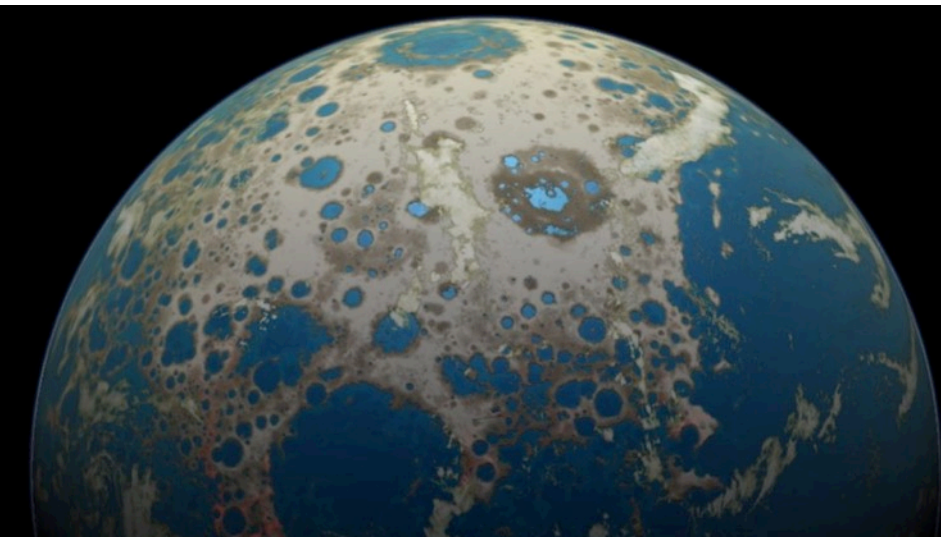
Artistic rendering of Hadean Earth (credit: Simone Marchi)



Heavily cratered lunar surface (credit: NASA)

Why study the moon?

- Moon is an unique archive of early Solar System history
 - Early terrestrial record erased by crustal recycling



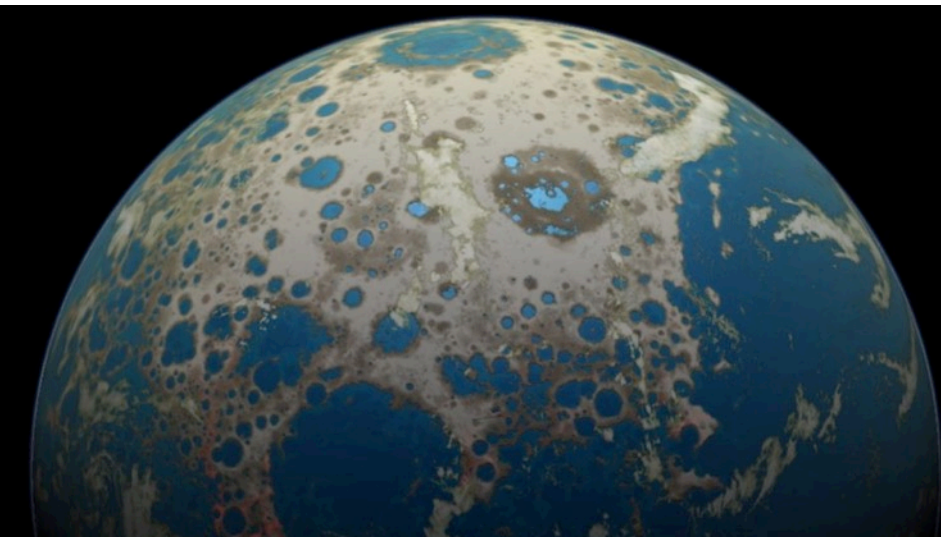
Artistic rendering of Hadean Earth (credit: Simone Marchi)



Heavily cratered lunar surface (credit: NASA)

Why study the moon?

- Moon is an unique archive of early Solar System history
 - Early terrestrial record erased by crustal recycling
 - Implications for emergence of life



Artistic rendering of Hadean Earth (credit: Simone Marchi)



Heavily cratered lunar surface (credit: NASA)

Why study the moon?

- Lunar history can be extrapolated to other solid surfaces in Solar System



Martian surface (credit: European Space Agency)

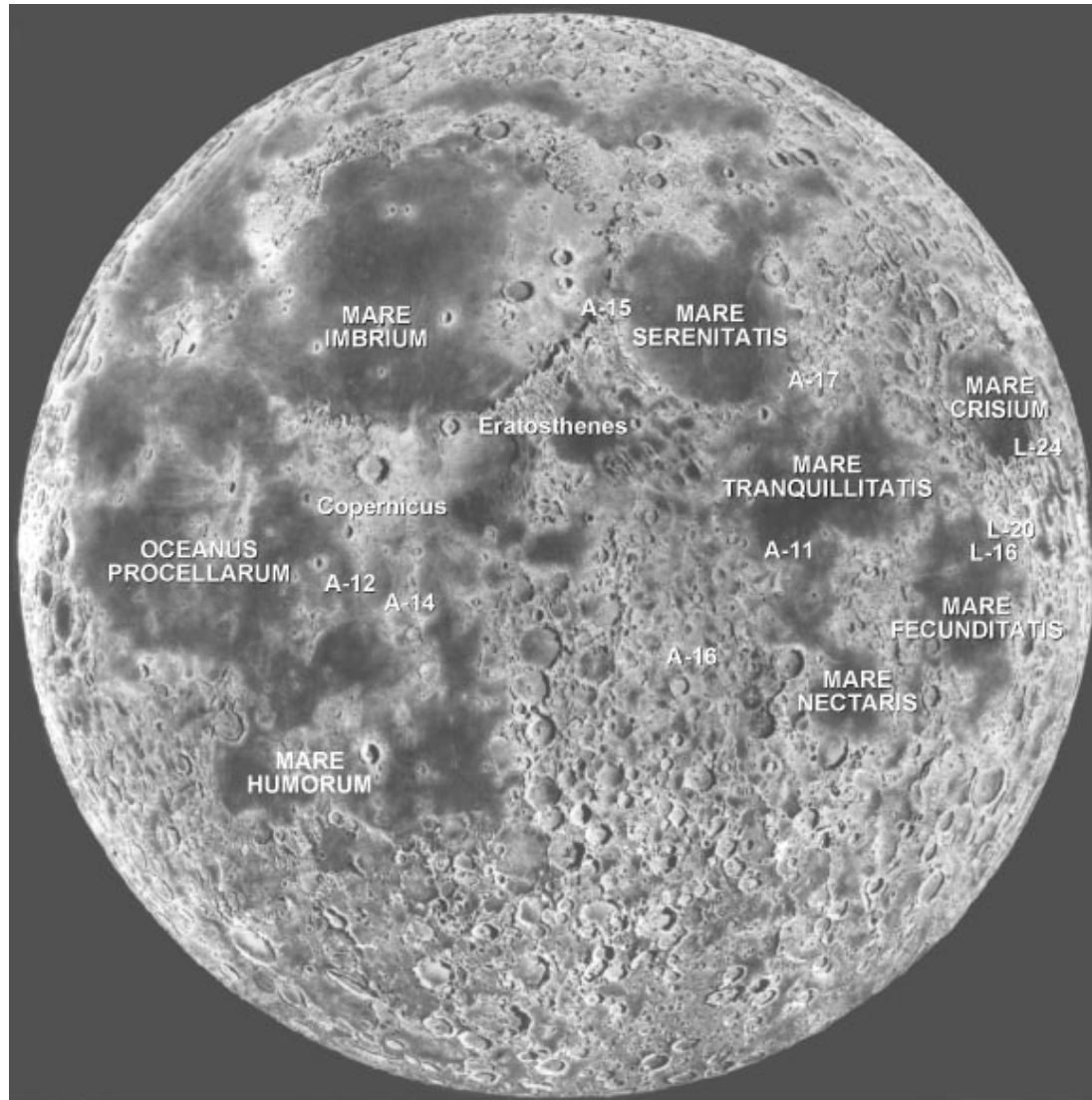
Why study the moon?

- Lunar history can be extrapolated to other solid surfaces in Solar System
 - Lunar chronology ties crater size frequency distributions to radiometric time

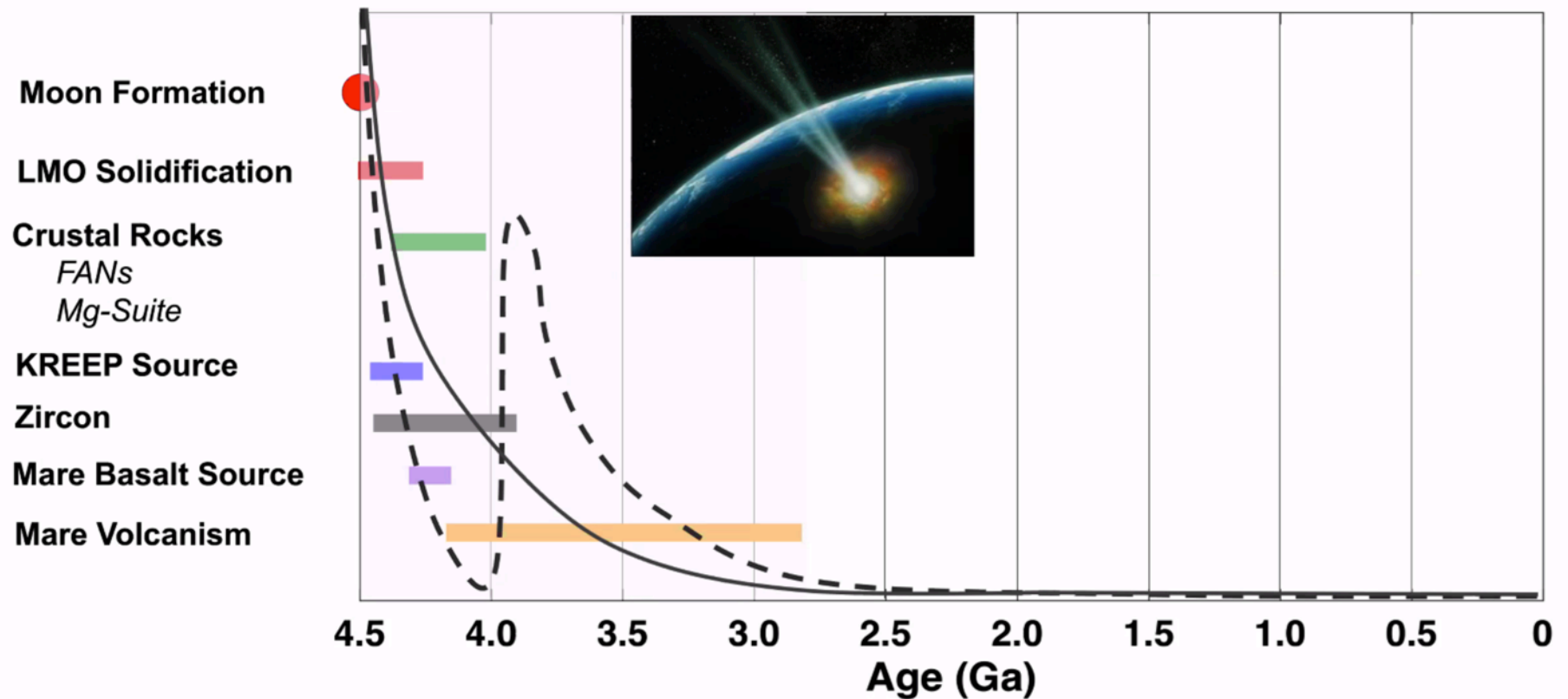


Martian surface (credit: European Space Agency)

The lunar surface



Lunar chronology



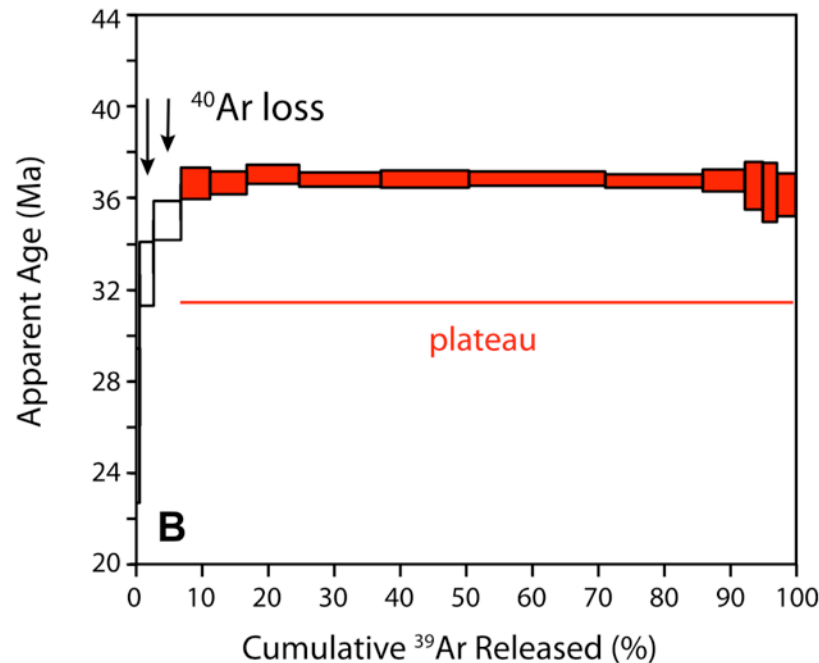
Credit: NASA

Argon geochronology

- ^{40}K ($t_{1/2} = 1.25 \text{ Ga}$) \rightarrow ^{40}Ca and ^{40}Ar
- Neutron irradiation produces ^{39}Ar from ^{40}K
- $^{40}\text{Ar}/^{39}\text{Ar}$ measured in unknown and compared to neutron fluence age monitor

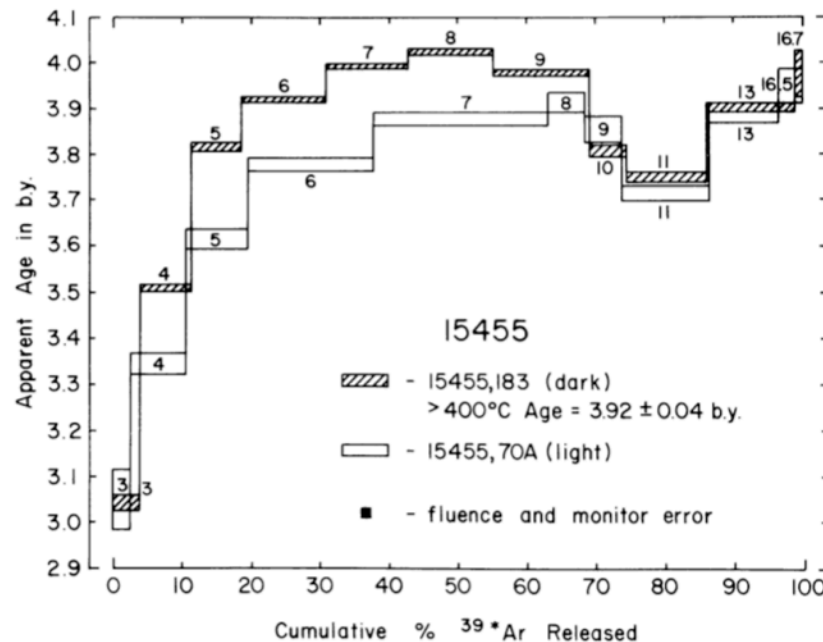
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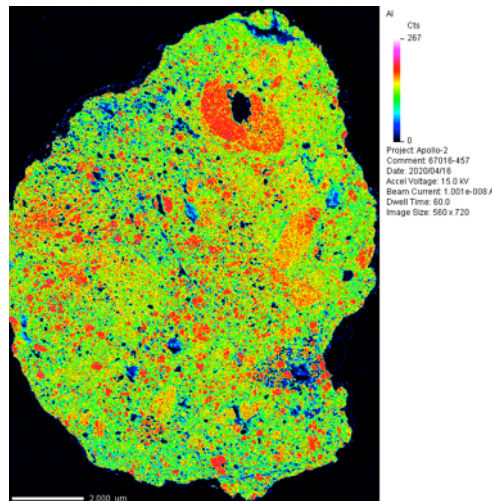
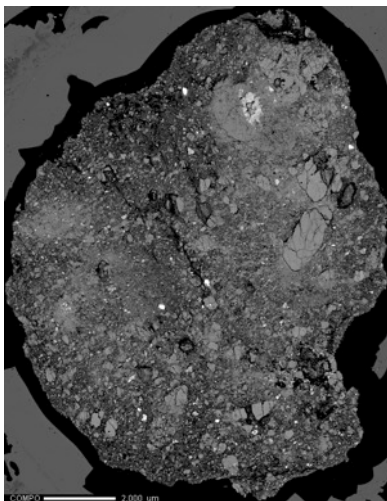


The UVLAMP $^{40}\text{Ar}/^{39}\text{Ar}$ method

- UV laser ablation microprobe (UVLAMP) analyses permit high-spatial resolution geochronologic investigations
 - Particularly useful for multi-generational materials
 - 193 nm wavelength, produces no collateral heating outside of ablation pit
 - Ancient nature of lunar materials require only tens of nanograms of ablated material
 - Sample targeting done in petrographic context

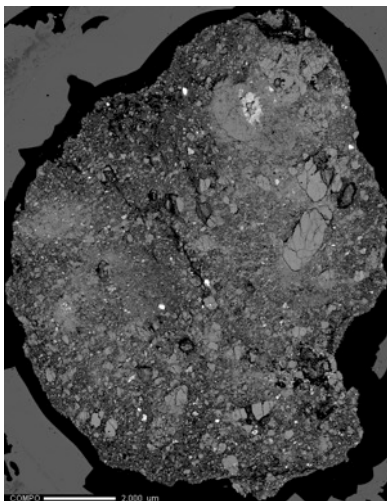
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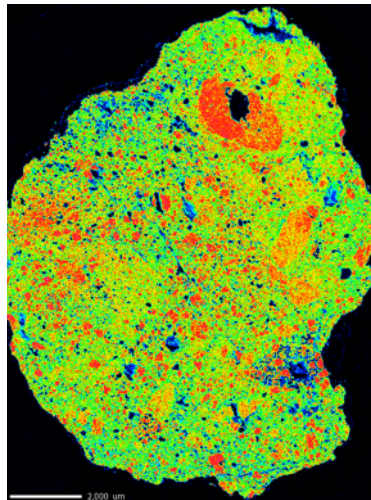


The UVLAMP $^{40}\text{Ar}/^{39}\text{Ar}$ method

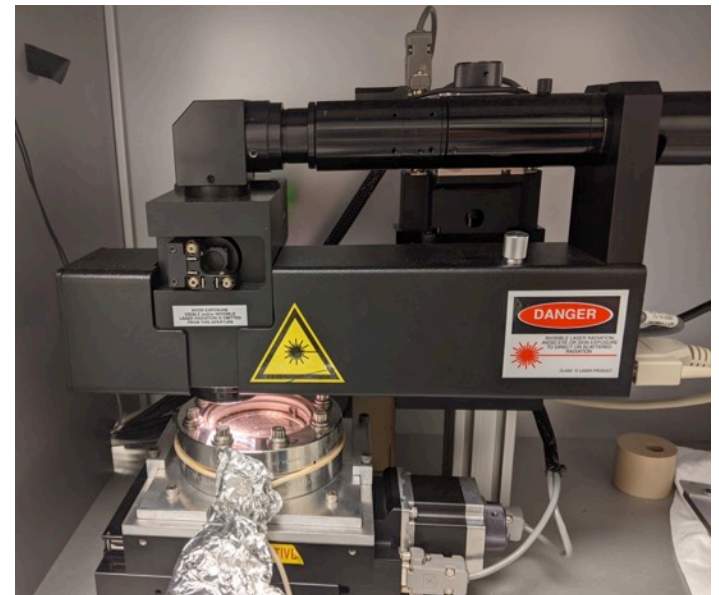
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COMPO
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Project Apollo-2
Comment: 67016-457
Date: 2020/04/16
Accel Voltage: 15.0 kV
Beam Current: 1.001e-008 A
Dwell Time: 60.0
Image Size: 560 x 720

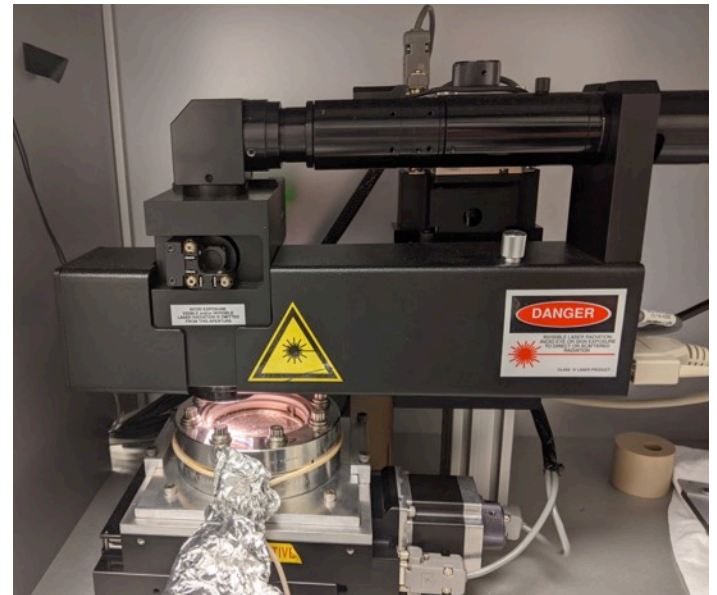
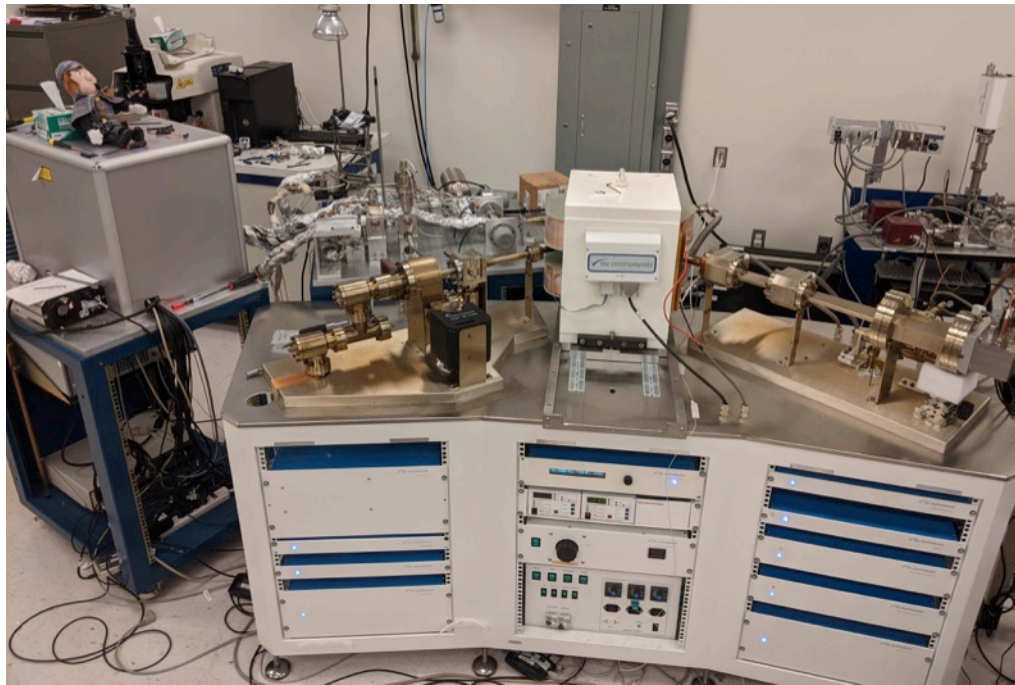


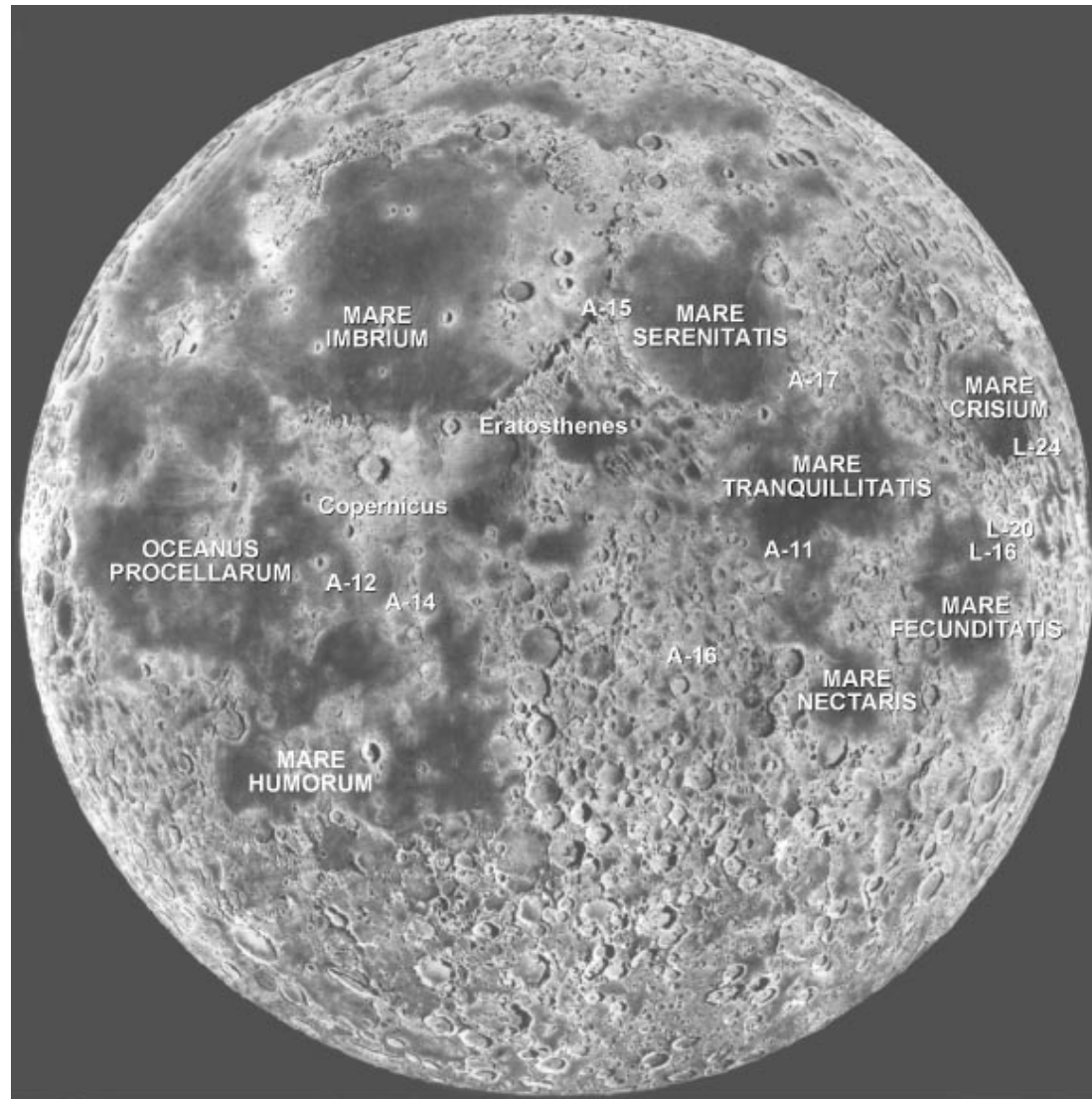
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Project Apollo-2
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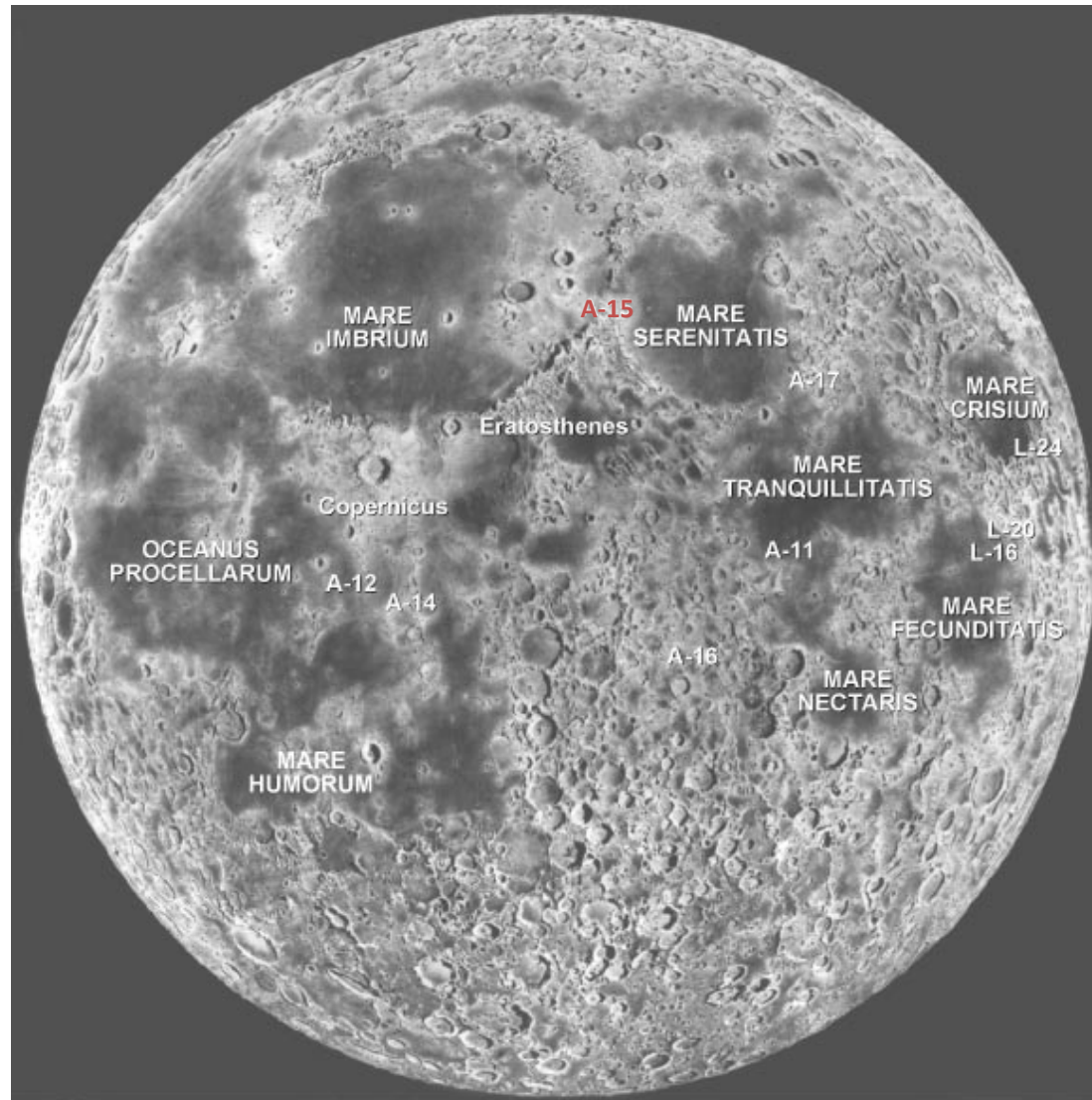


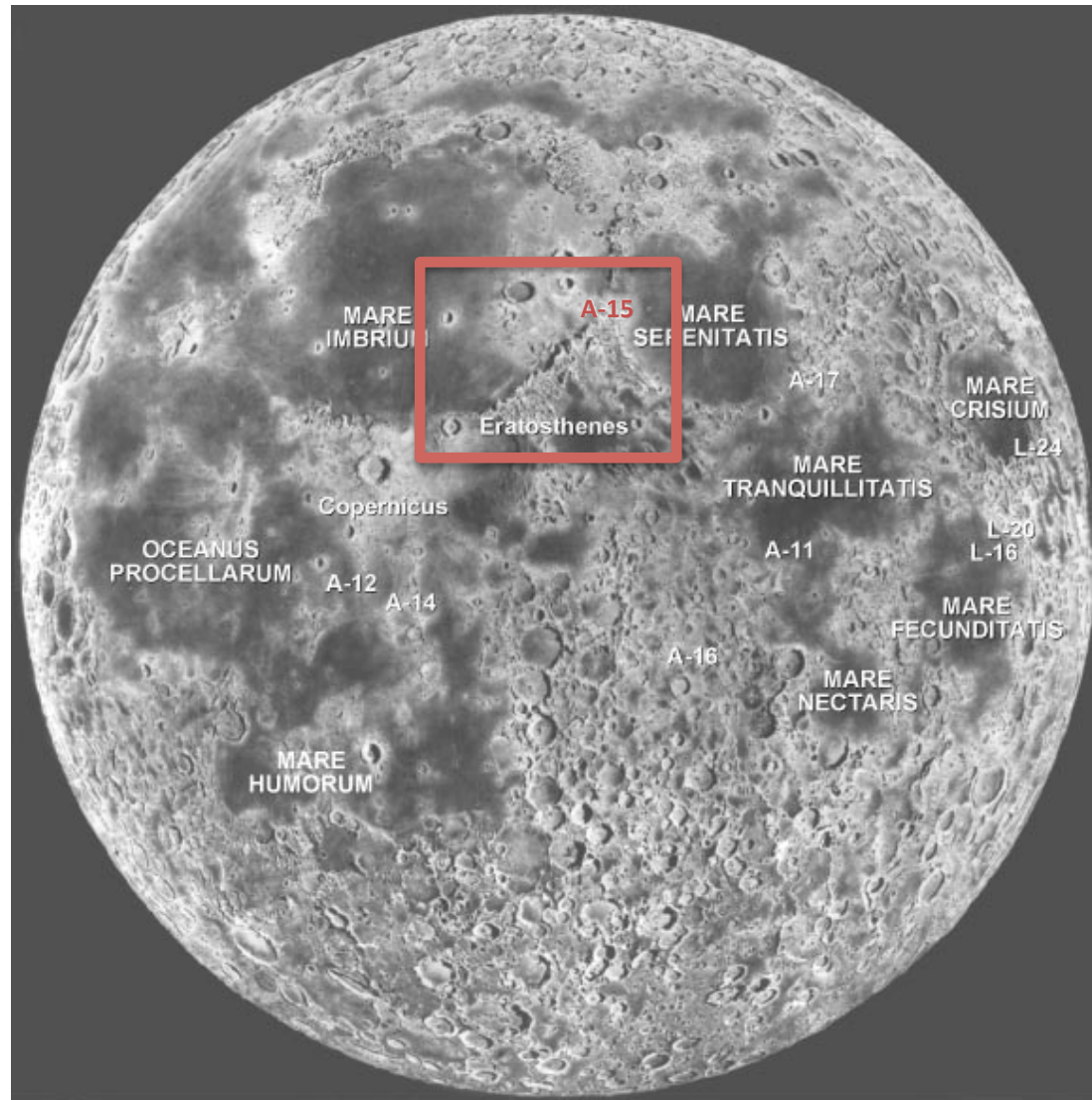
The UVLAMP $^{40}\text{Ar}/^{39}\text{Ar}$ method

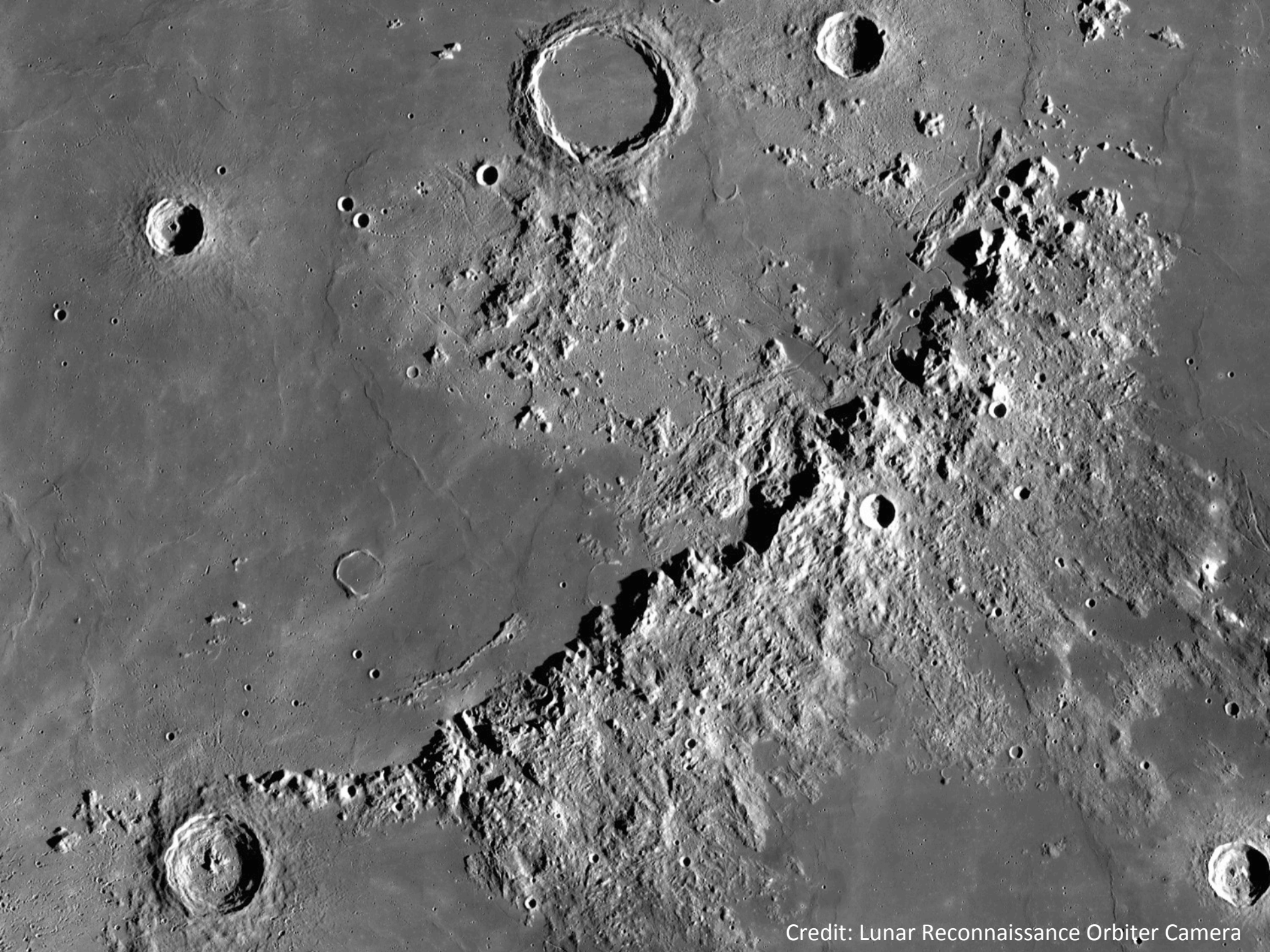
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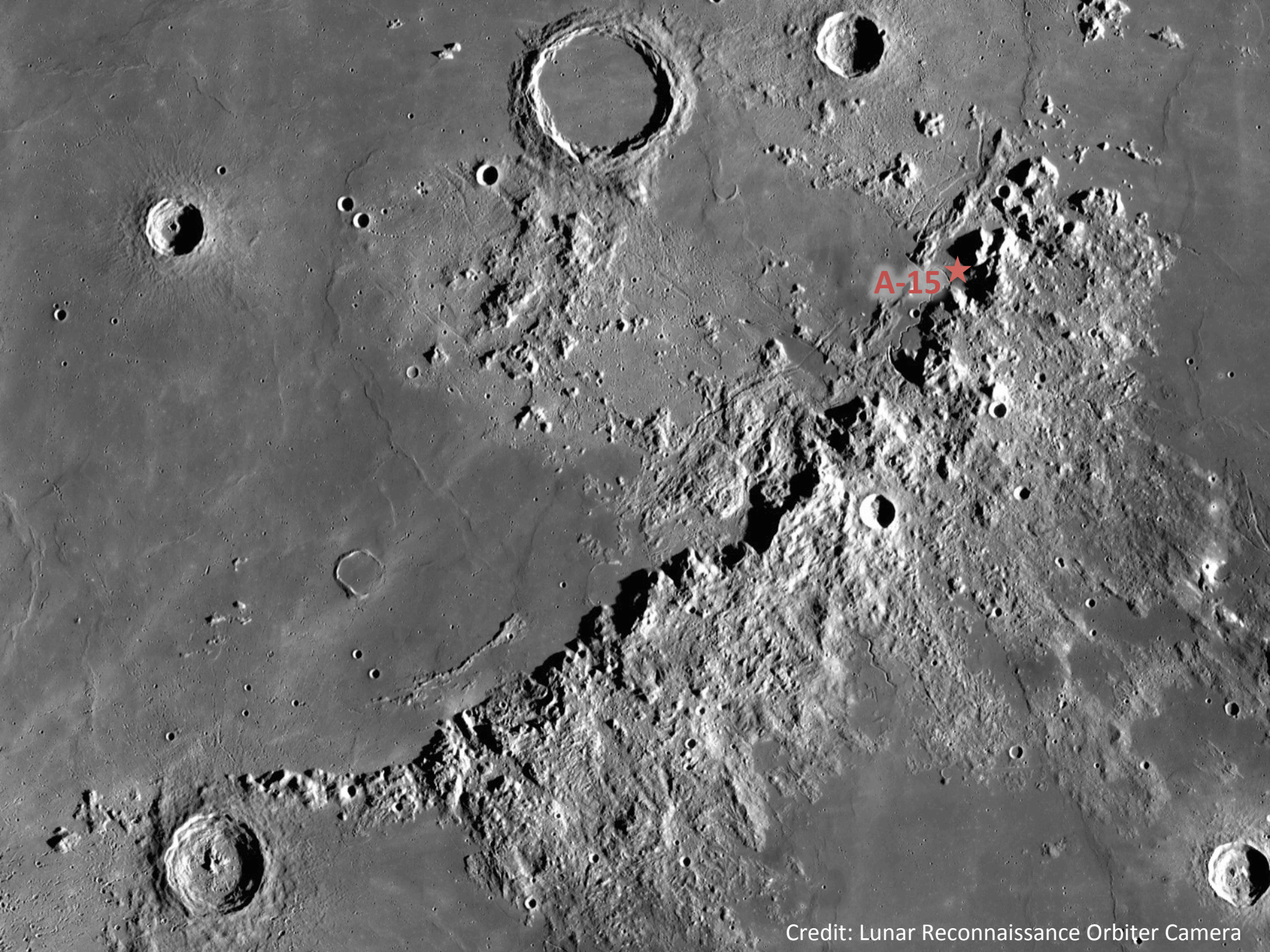




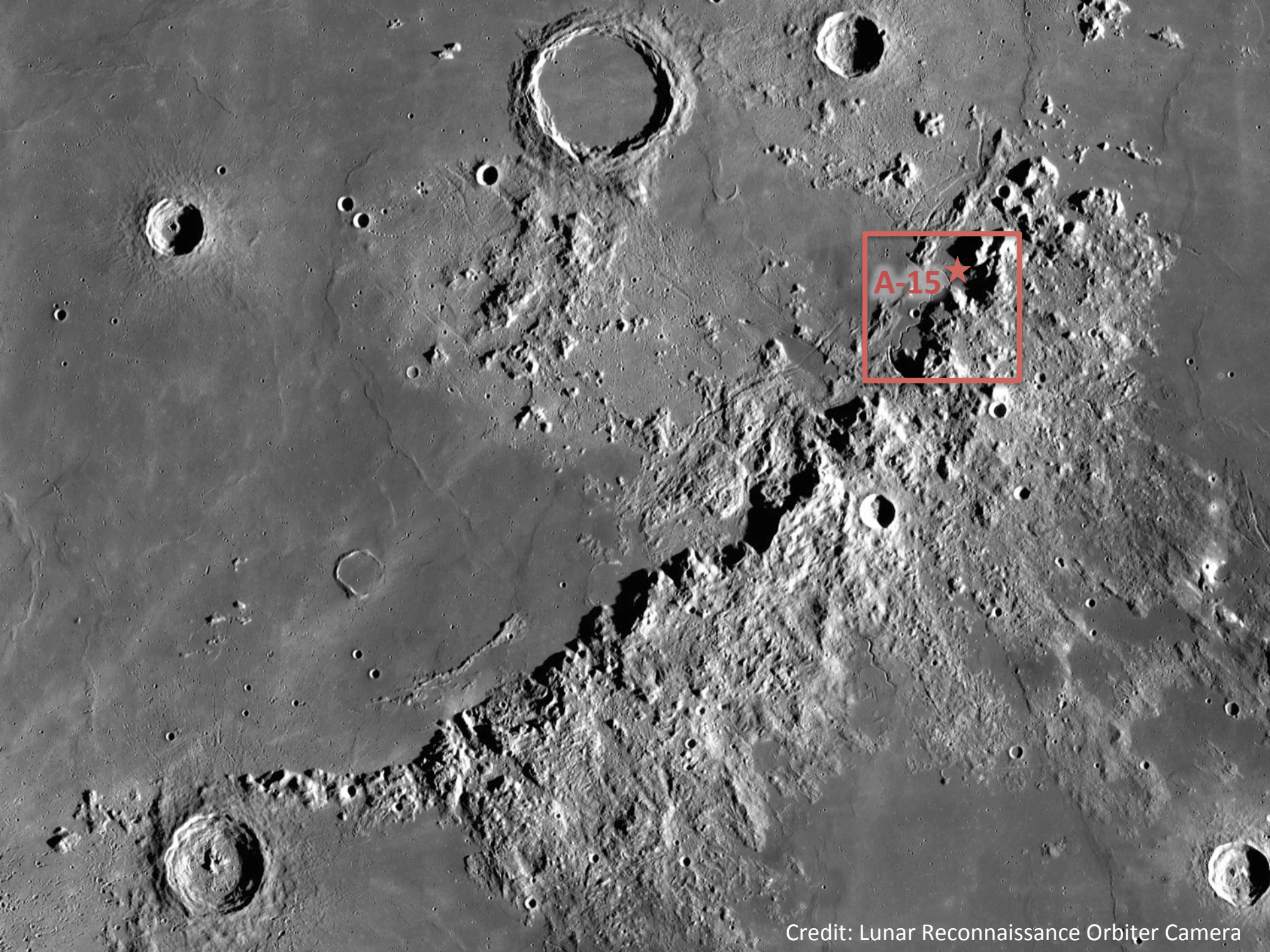




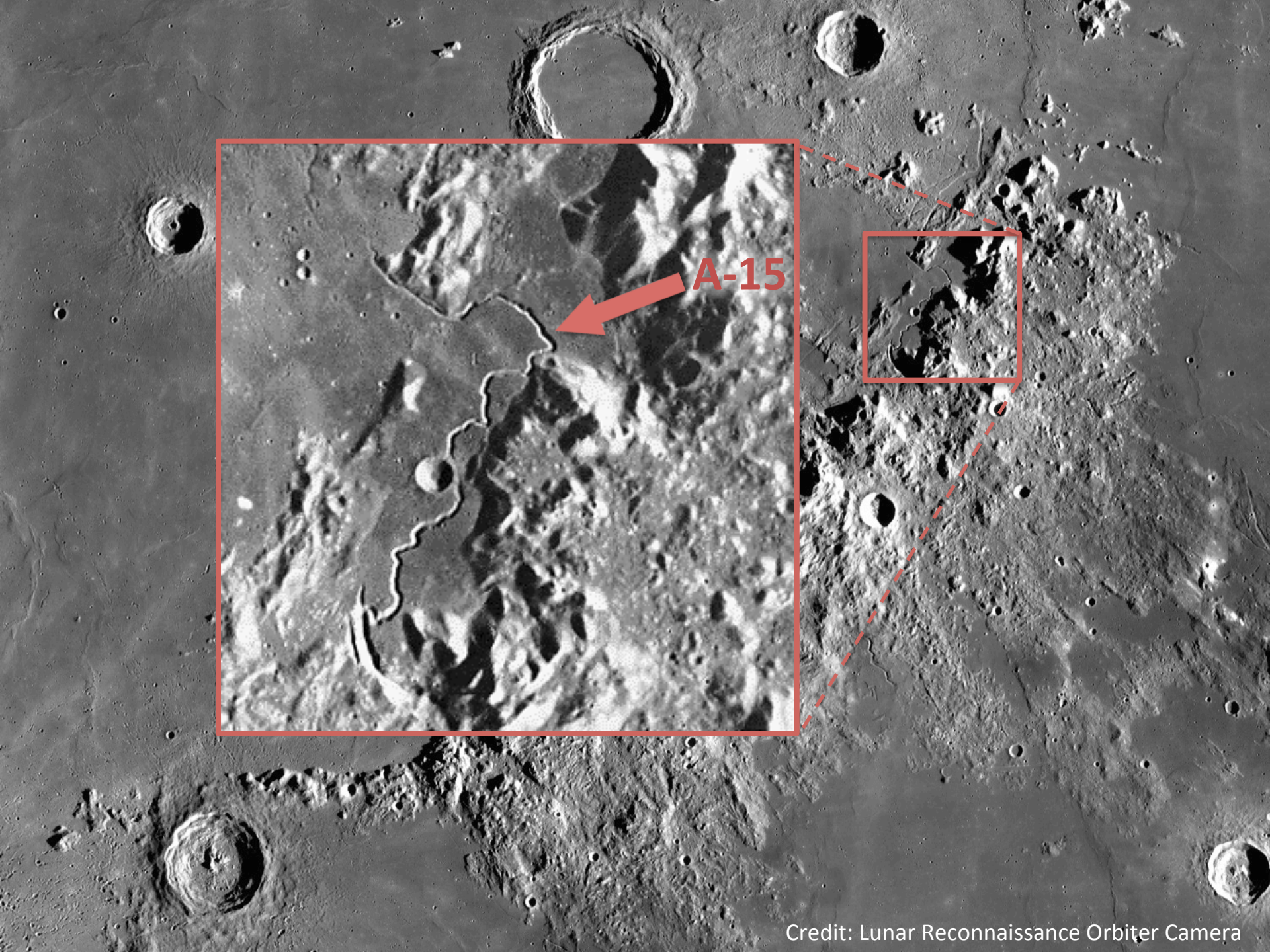
Credit: Lunar Reconnaissance Orbiter Camera



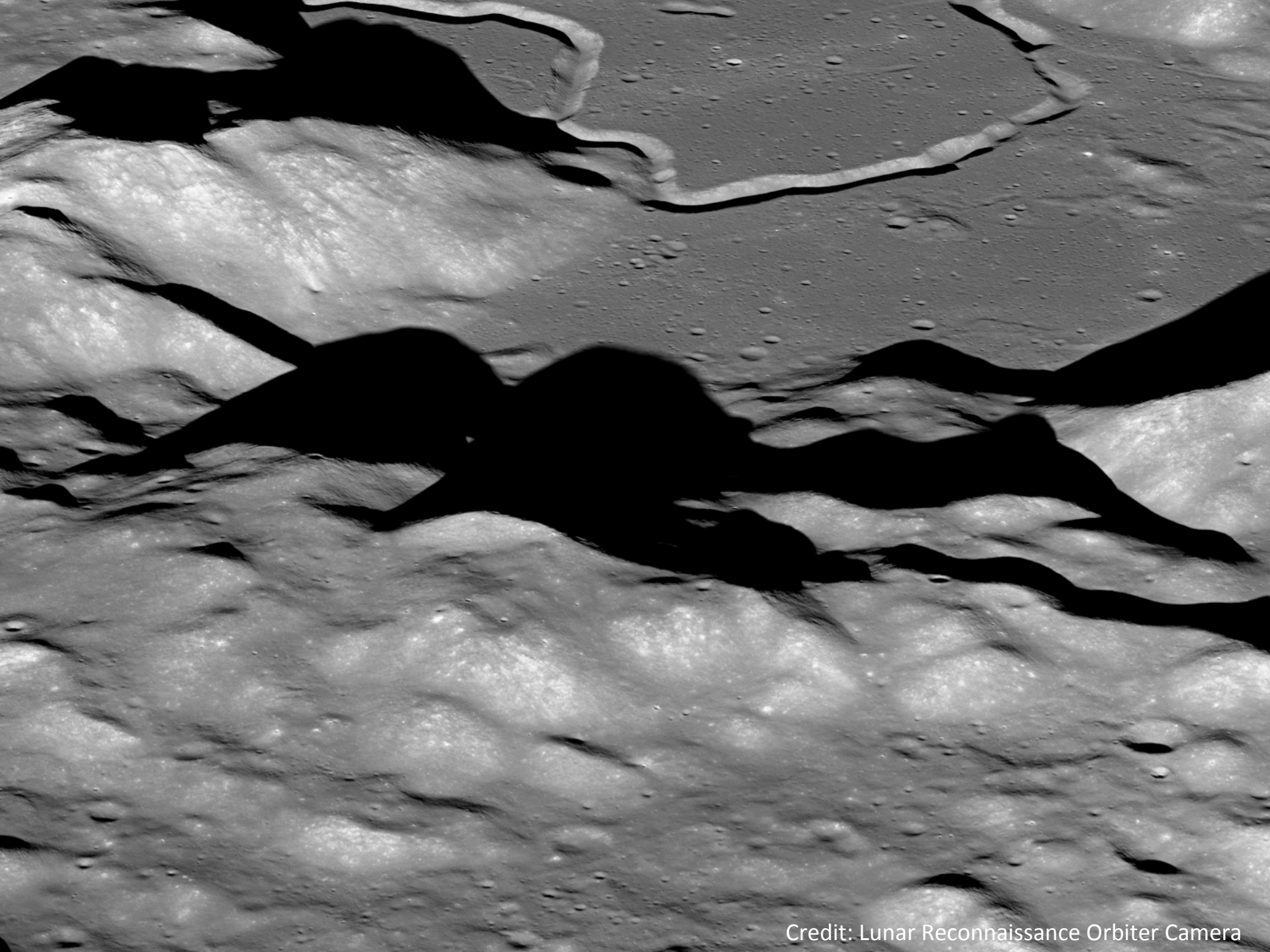
Credit: Lunar Reconnaissance Orbiter Camera



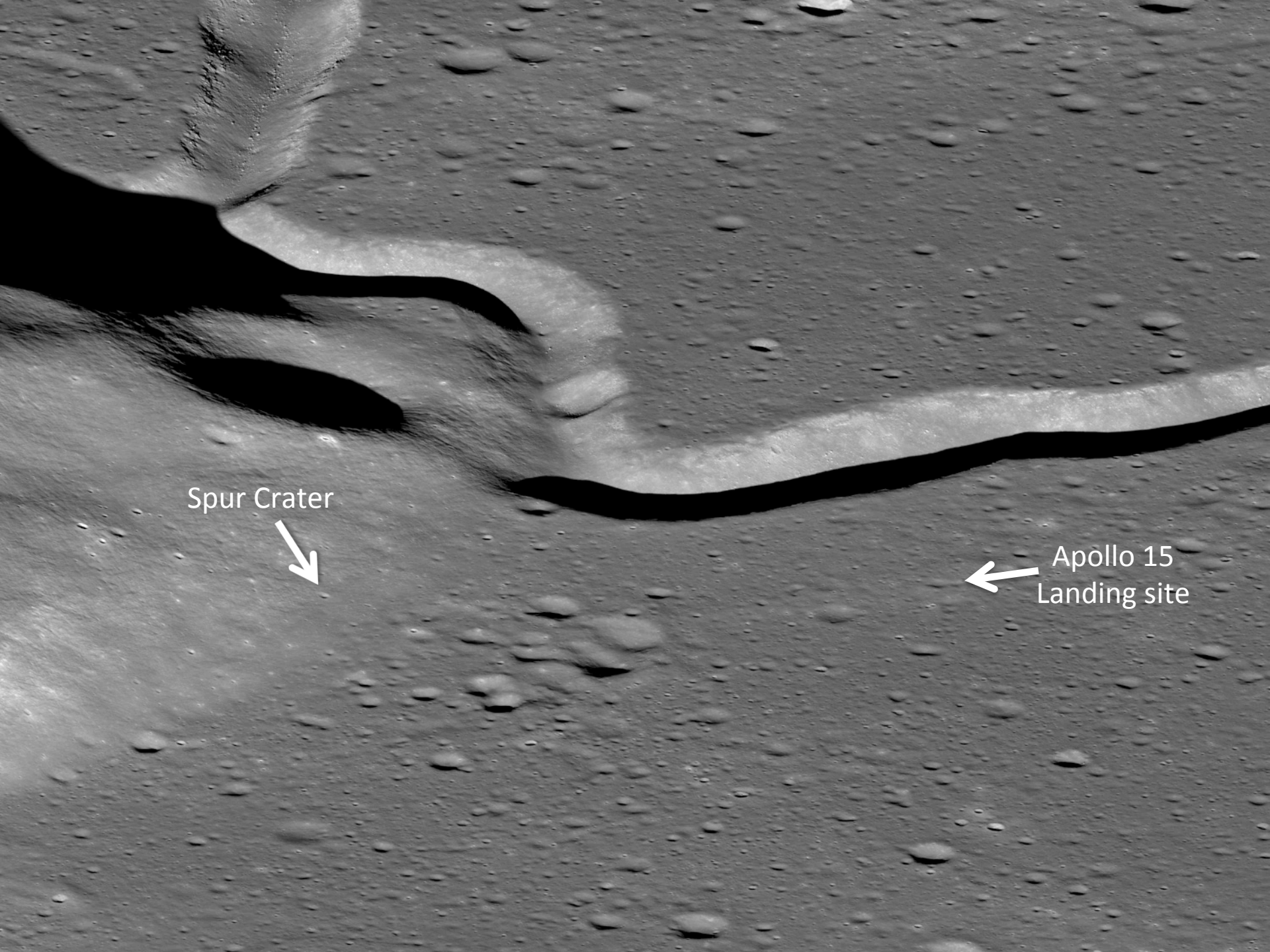
Credit: Lunar Reconnaissance Orbiter Camera



Credit: Lunar Reconnaissance Orbiter Camera



Credit: Lunar Reconnaissance Orbiter Camera



Spur Crater

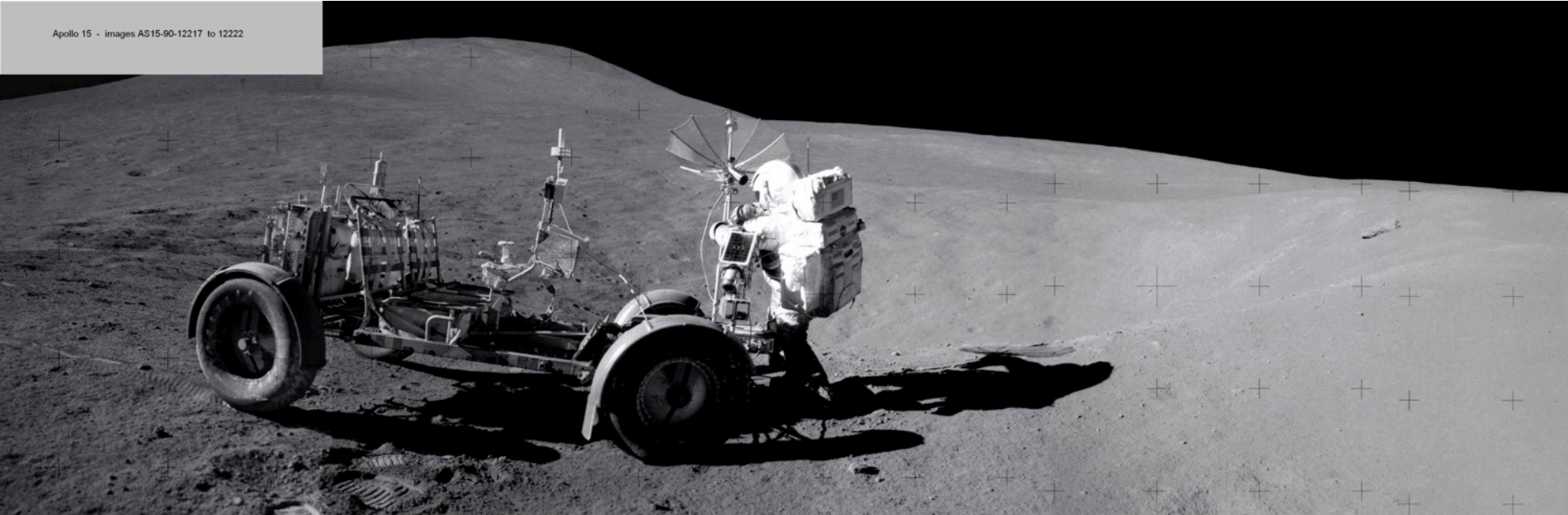


Apollo 15
Landing site



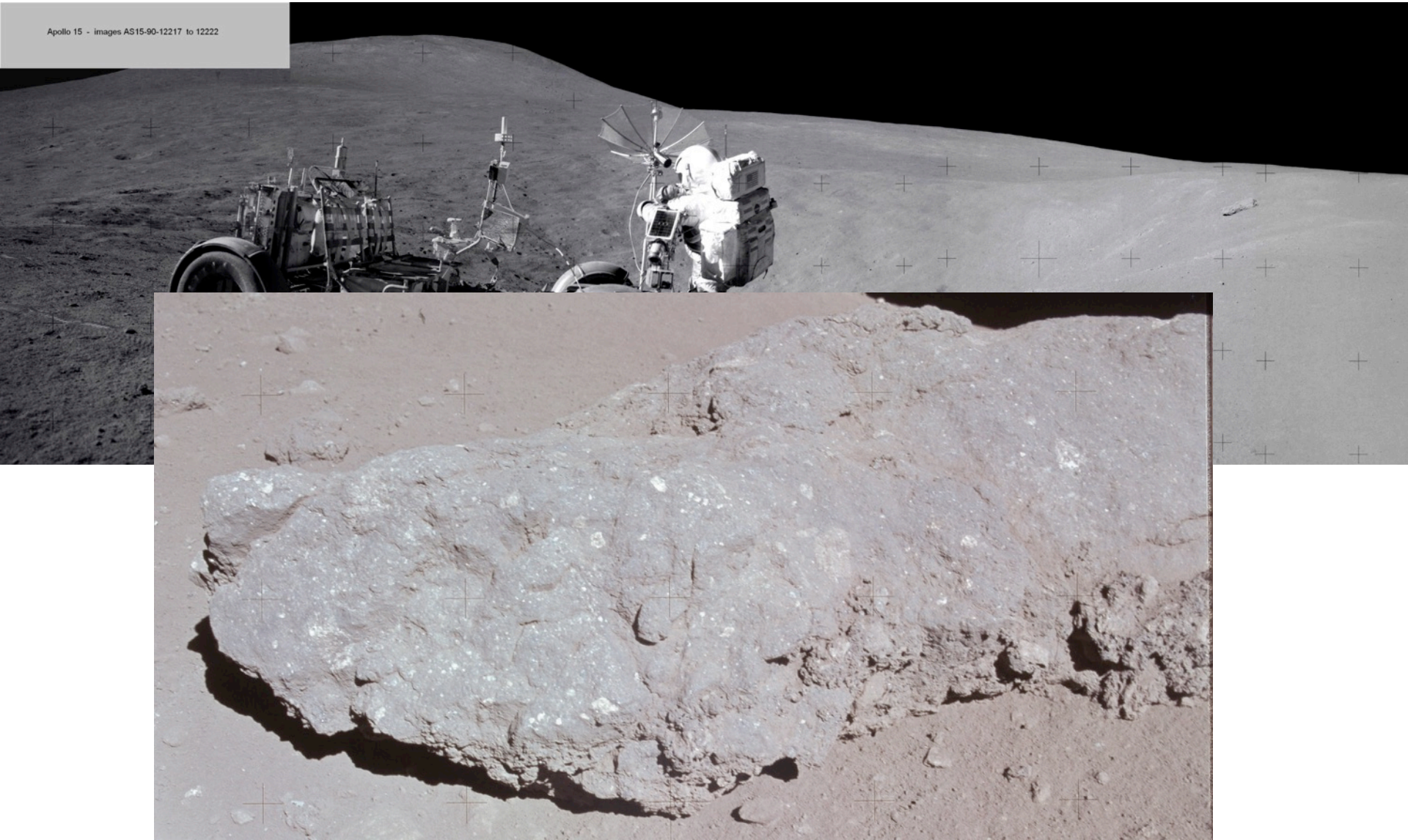
Spur Crater – 15455 and 15445

Apollo 15 - Images AS15-90-12217 to 12222



Spur Crater – 15455 and 15445

Apollo 15 - Images AS15-90-12217 to 12222



The Apollo 15 'white an black' rocks



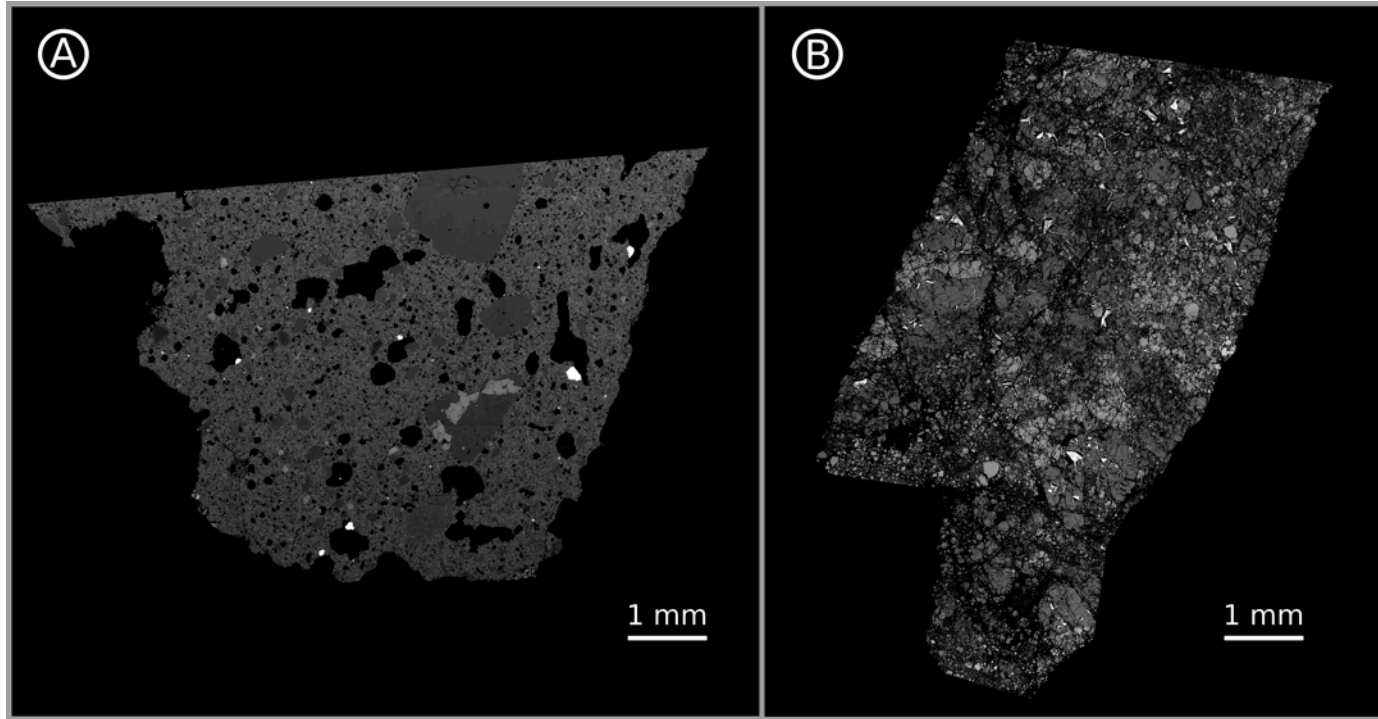
15445



15455

- Macroscopic fragments of plutonic rock hosted within a fragment-laden impact melt breccia
- The two samples are geochemically and petrographically linked

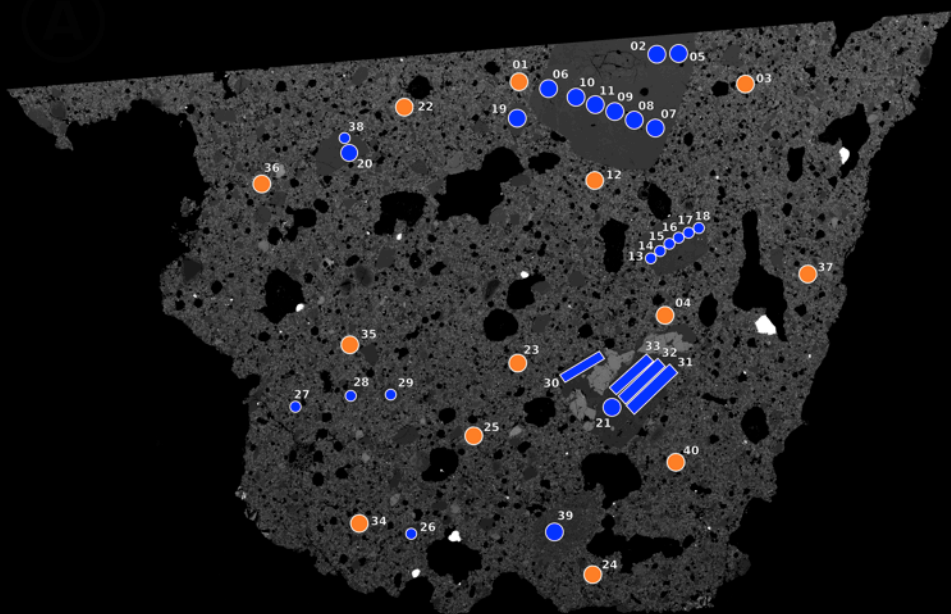
The Apollo 15 'white an black' rocks



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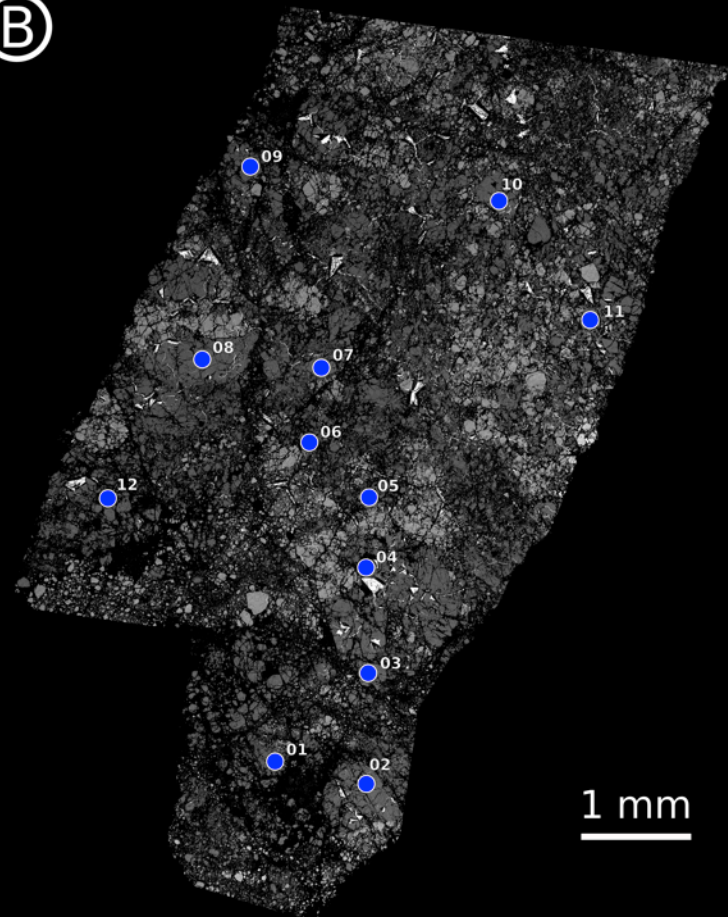
15455,383 and 15455,386

(A)



1 mm

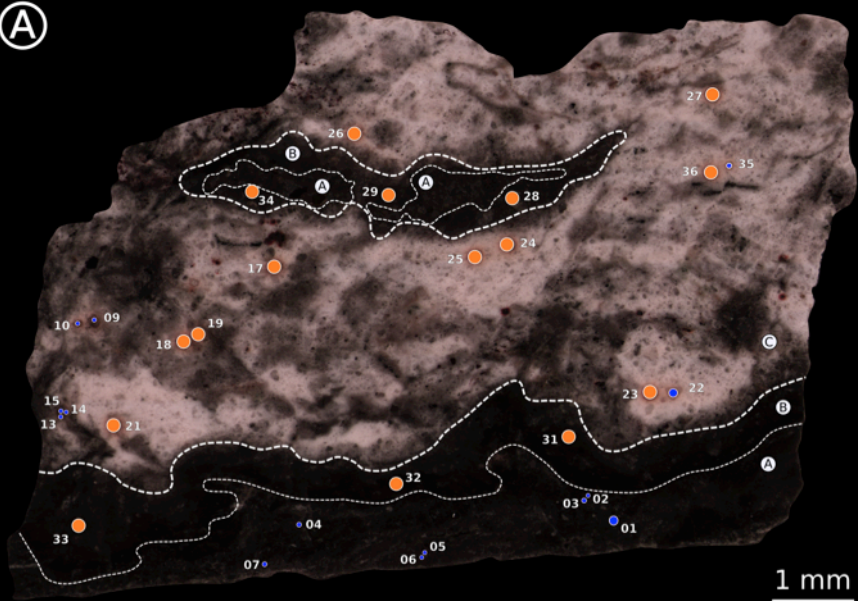
(B)



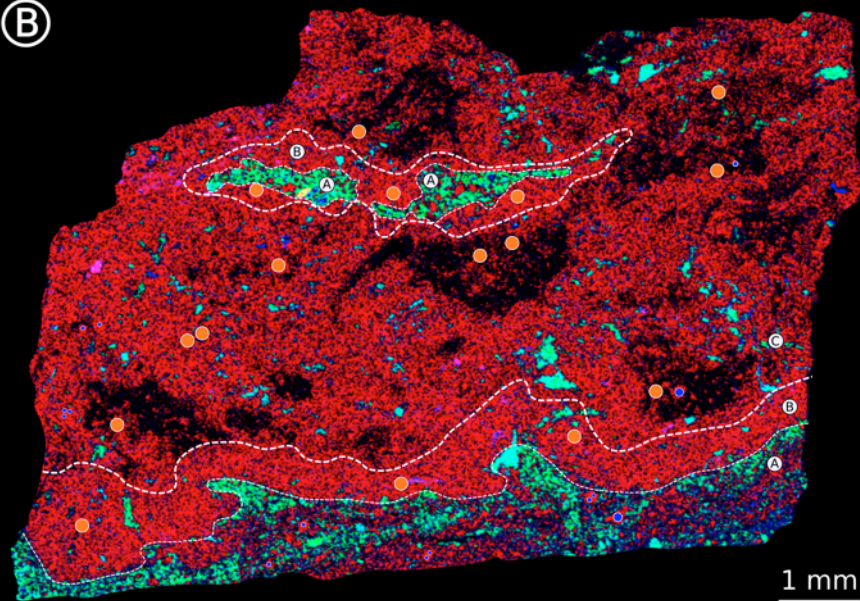
1 mm

15455,405

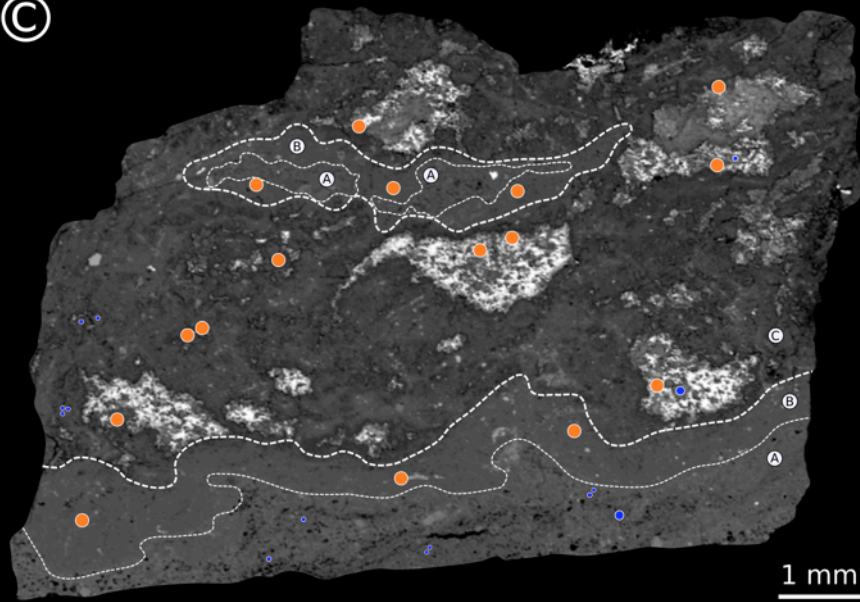
Ⓐ



Ⓑ

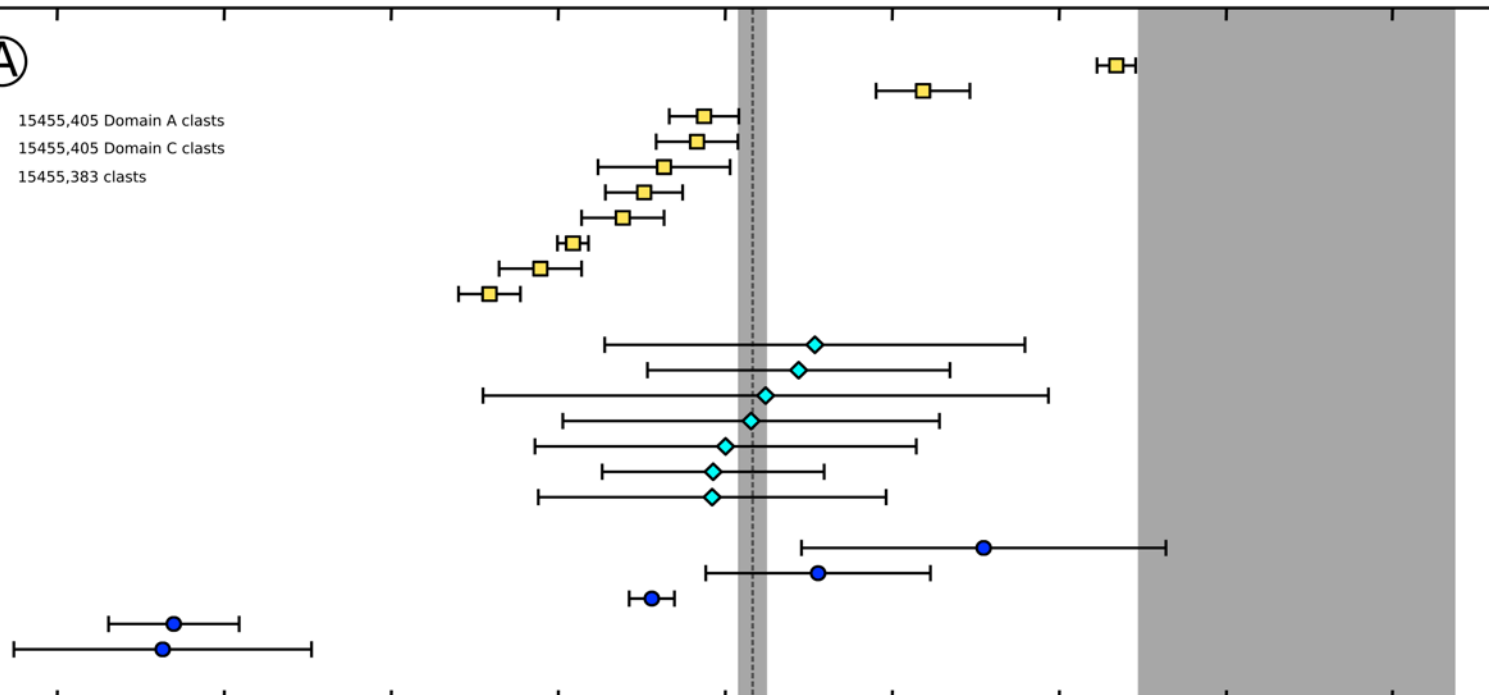


Ⓒ



(A)

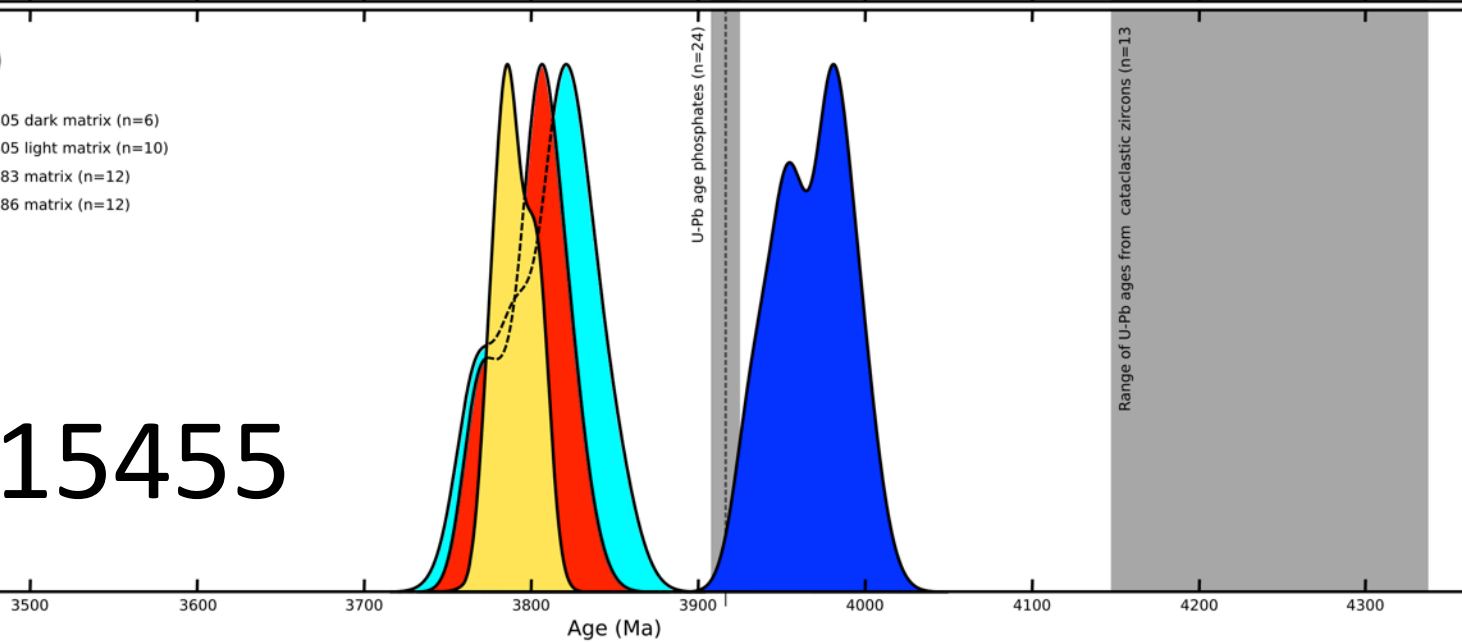
- 15455,405 Domain A clasts
- ◆ 15455,405 Domain C clasts
- 15455,383 clasts

**(B)**

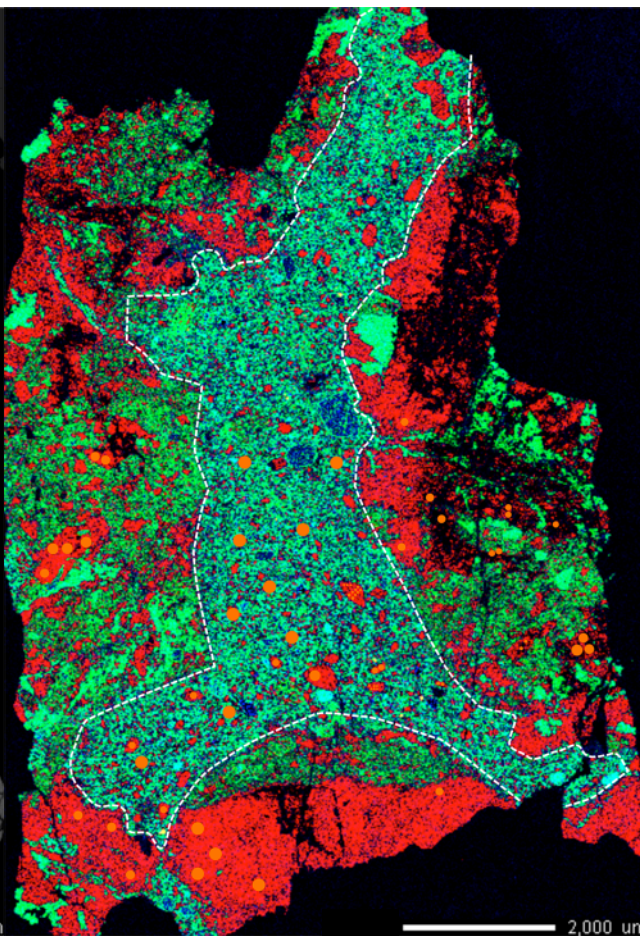
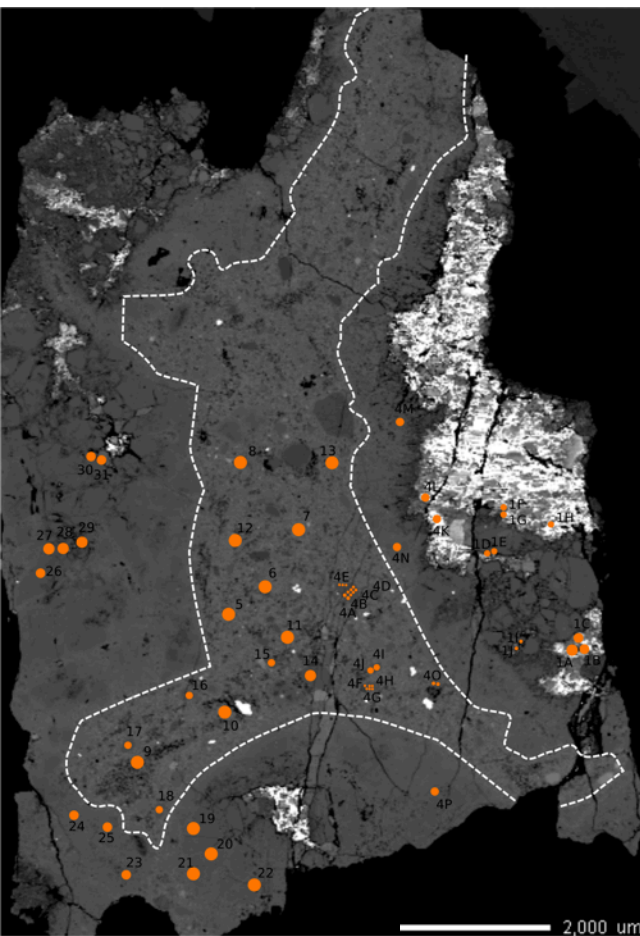
- 405 dark matrix (n=6)
- 405 light matrix (n=10)
- 383 matrix (n=12)
- 386 matrix (n=12)

Relative Probability

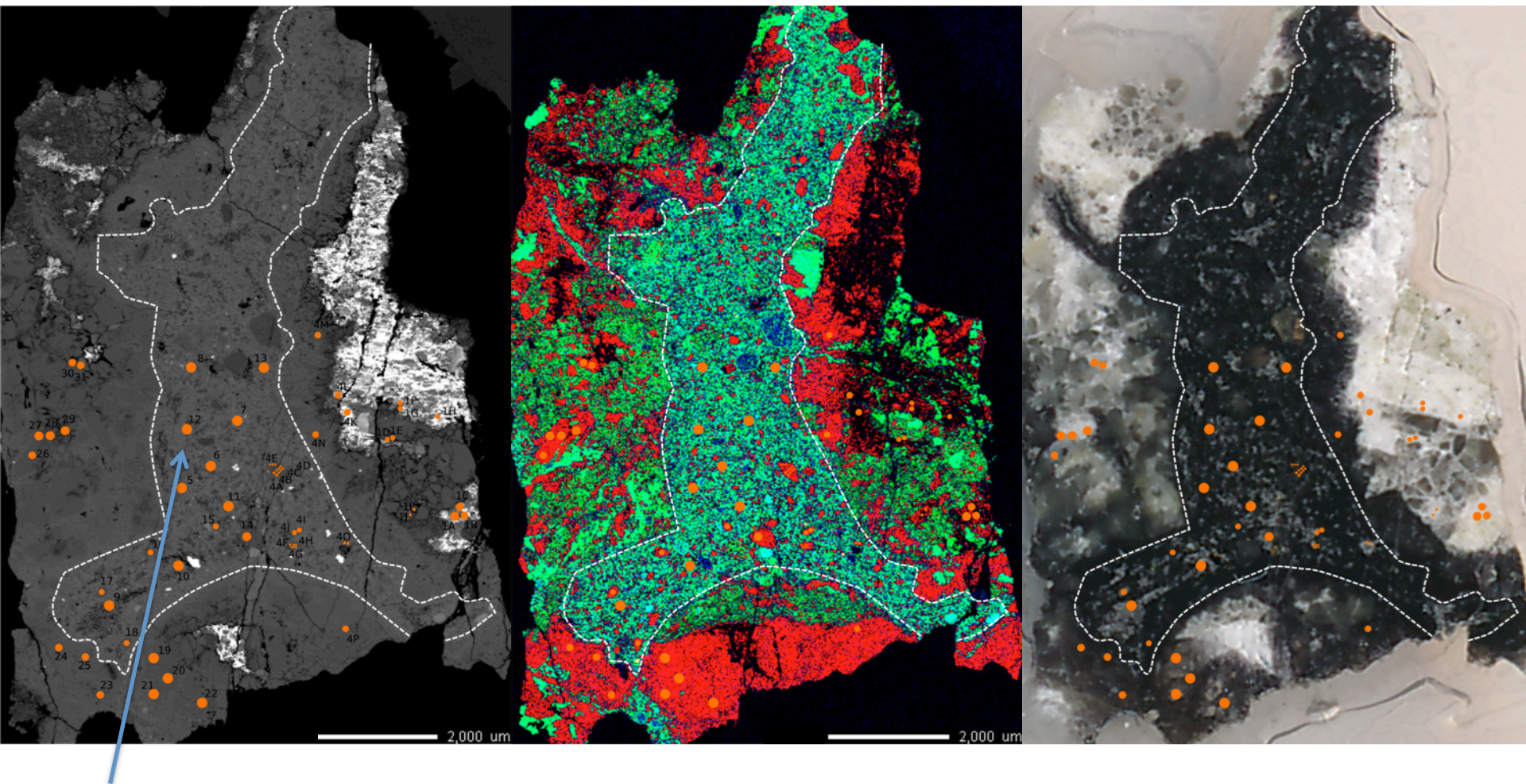
15455



15445

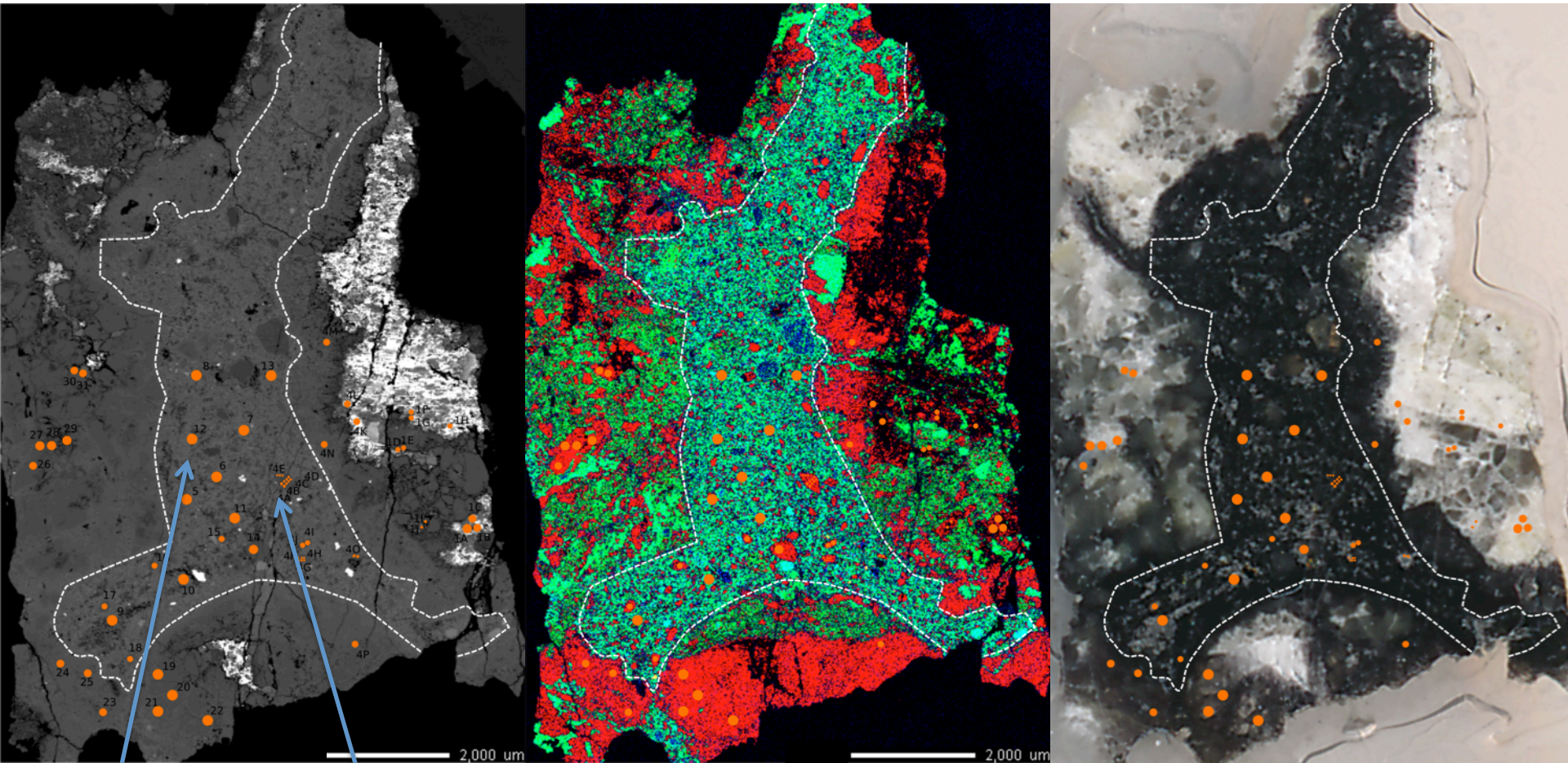


15445



$3668 \pm 33 \text{ Ma}$
($n=8$)

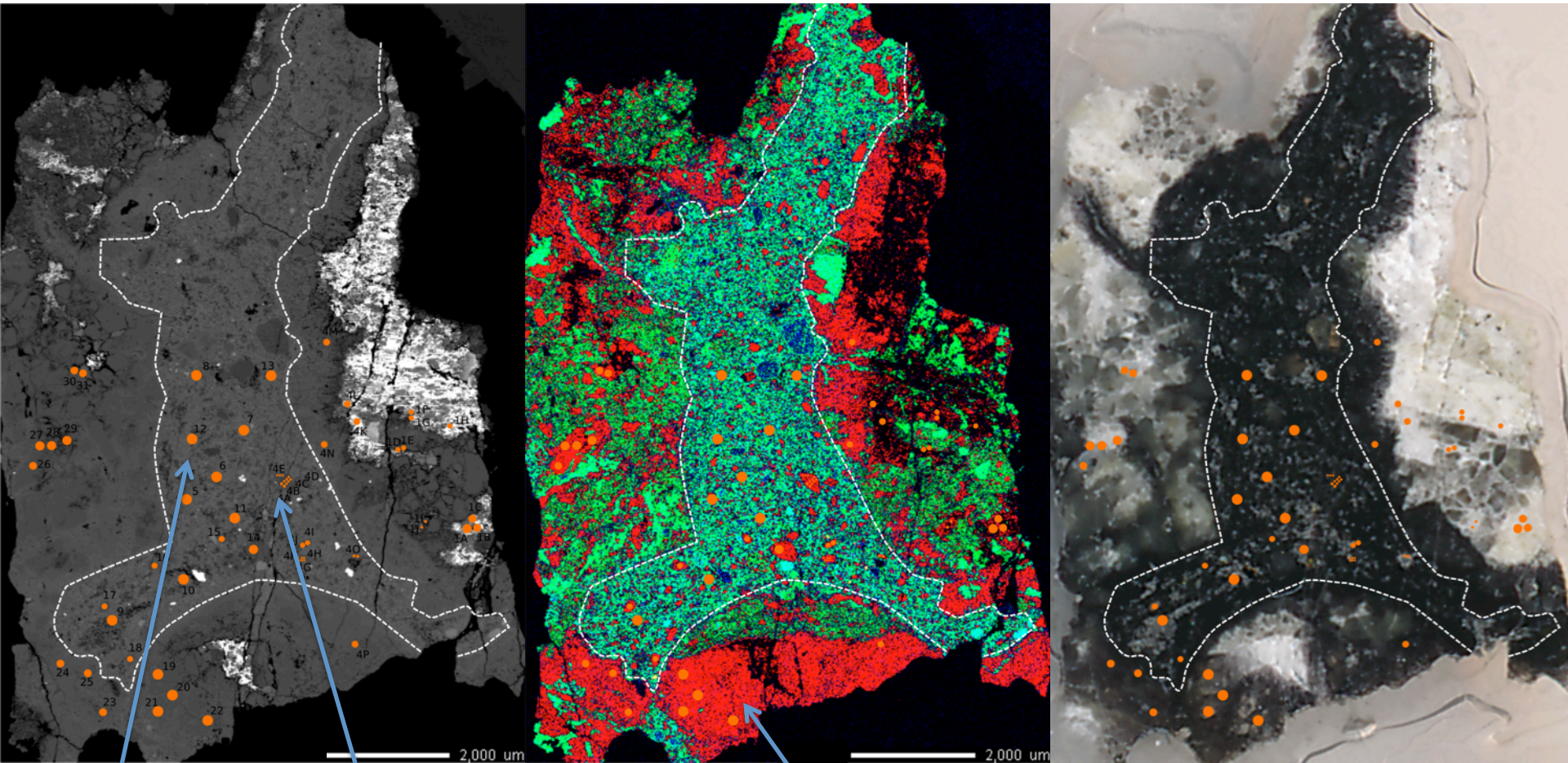
15445



3654 ± 86 Ma
to 4078 ± 43 Ma

3668 ± 33 Ma
(n=8)

15445

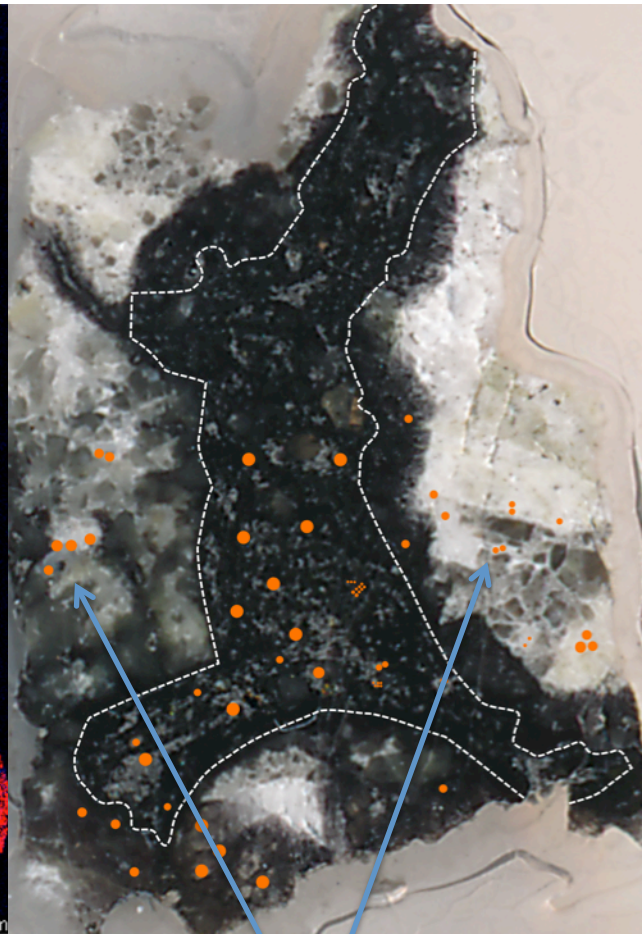
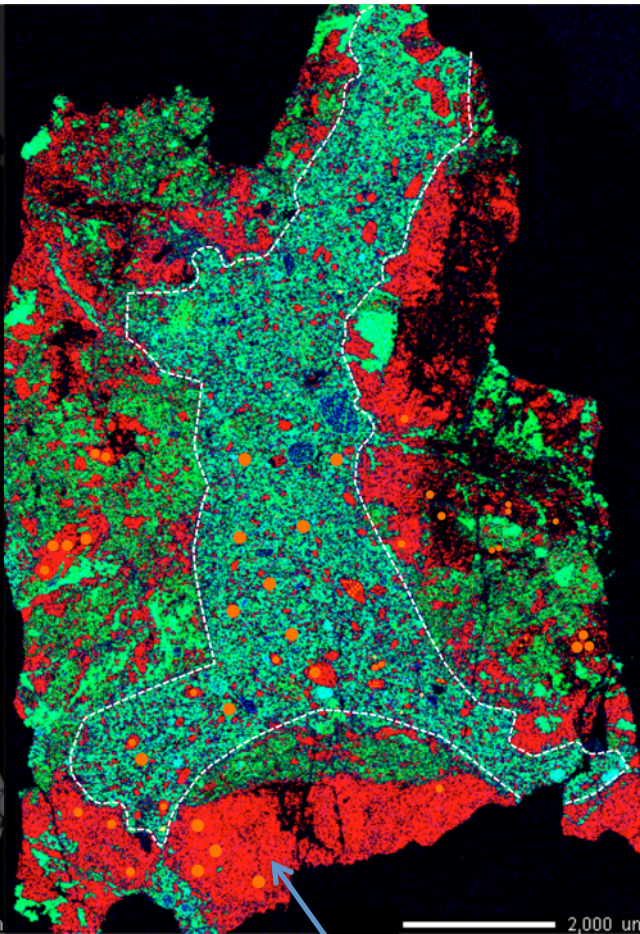
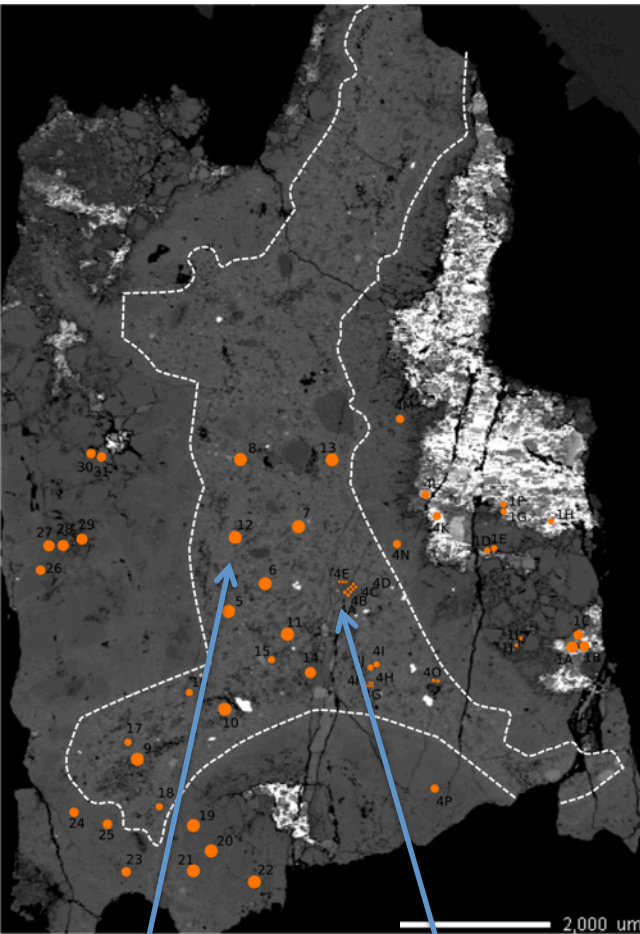


3668 ± 33 Ma
(n=8)

3654 ± 86 Ma
to 4078 ± 43 Ma

3922 ± 23 Ma (n=9)

15445



$3668 \pm 33 \text{ Ma}$
($n=8$)

$3654 \pm 86 \text{ Ma}$
to $4078 \pm 43 \text{ Ma}$

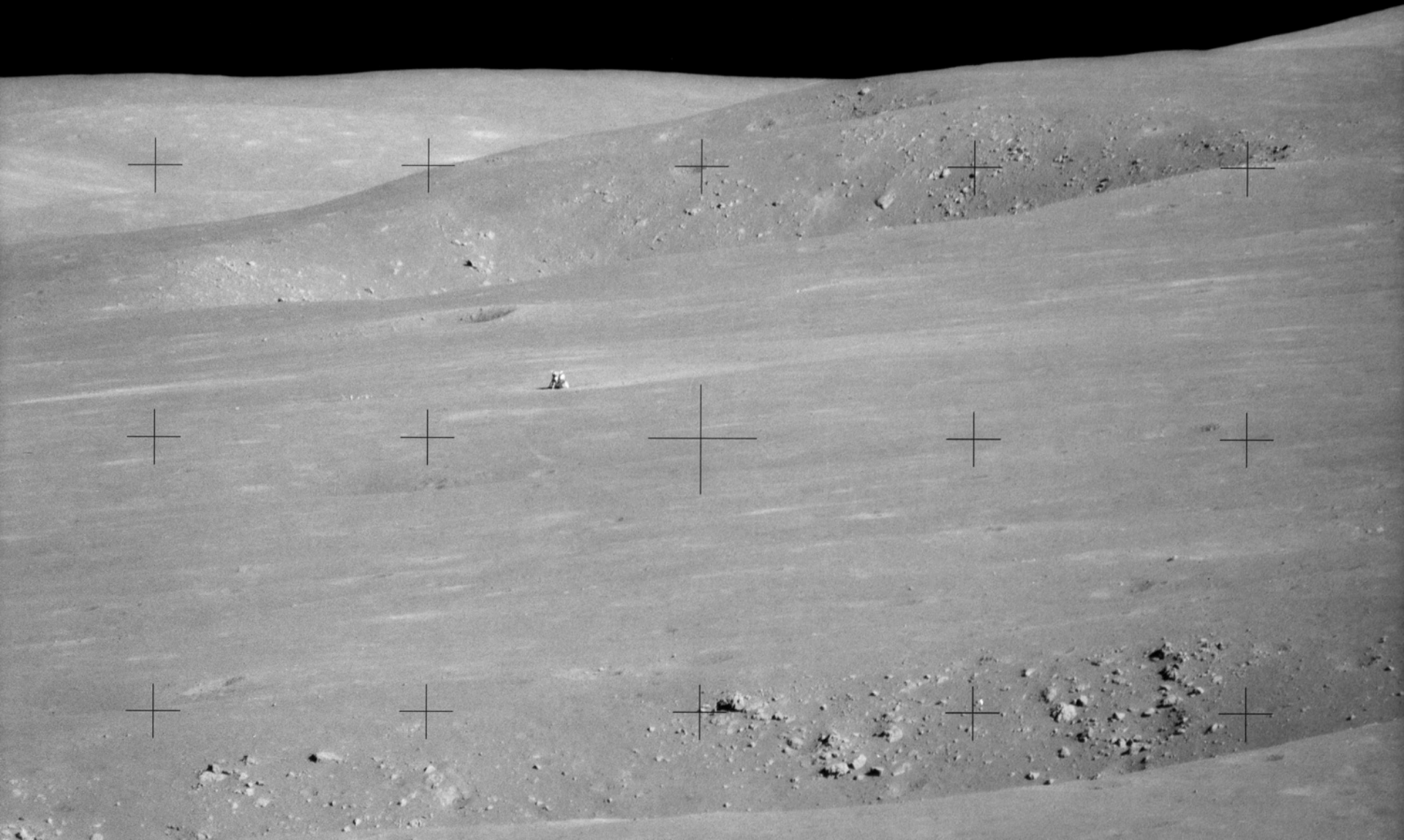
$3922 \pm 23 \text{ Ma}$ ($n=9$)

$3738 \pm 29 \text{ Ma}$
to $4028 \pm 26 \text{ Ma}$

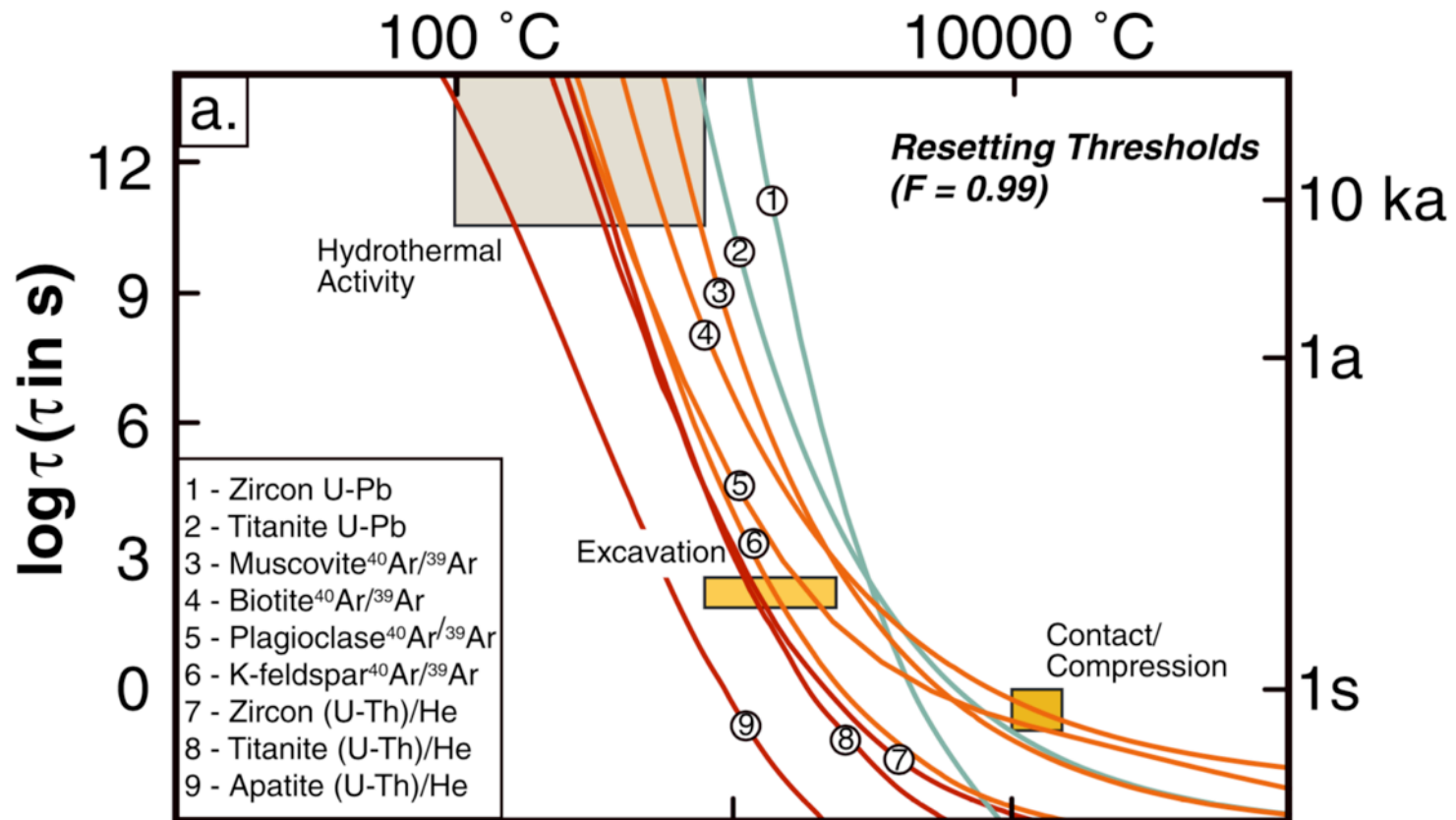
Conclusions

- High-spatial resolution UVLAMP $^{40}\text{Ar}/^{39}\text{Ar}$ analyses undertaken on the 'white and black' Apollo 15 samples
 - Protracted history of impact bombardment recorded spanning ca. 500 Ma
 - Impact melt generation in 15445 ca. 3670 Ma but ca. 3800 Ma in 15455
 - Despite geologic expectation, these samples may not record impact melt generation associated with the Imbrium basin forming impact
 - Powerful tool, especially when used as a complement to other high-precision chronometers such as U-Pb

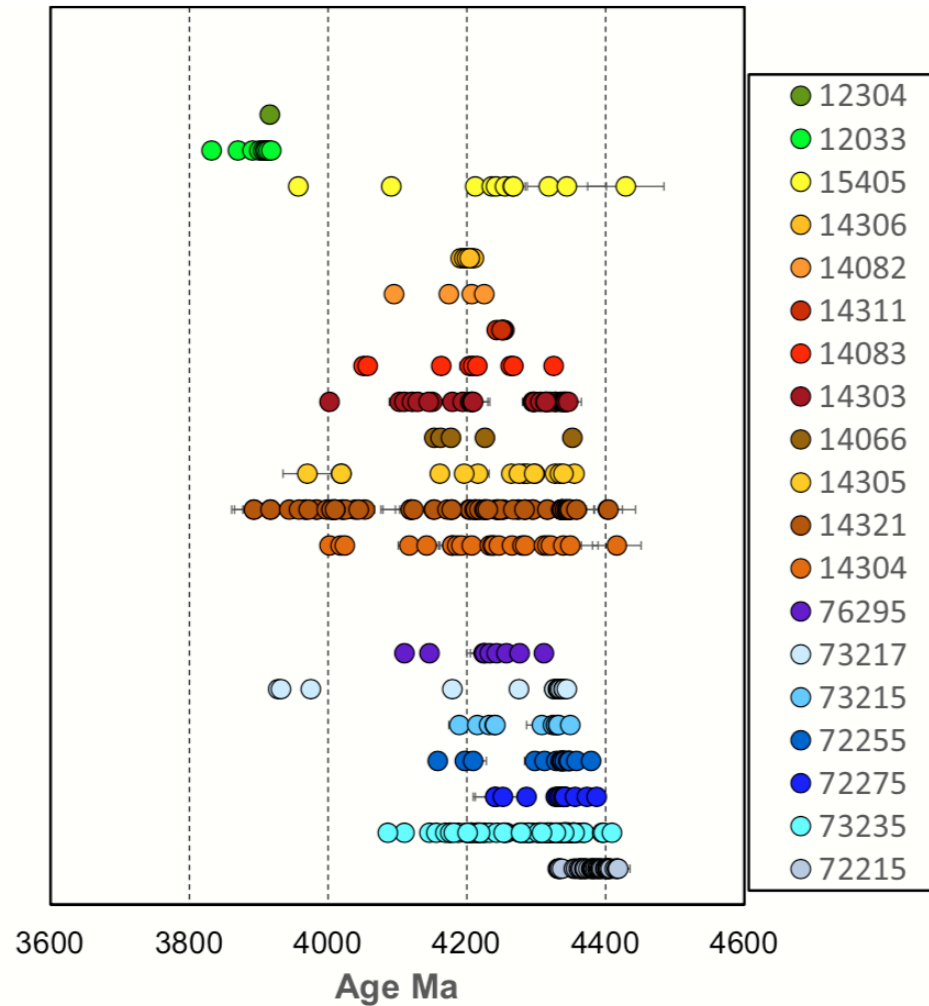
Questions



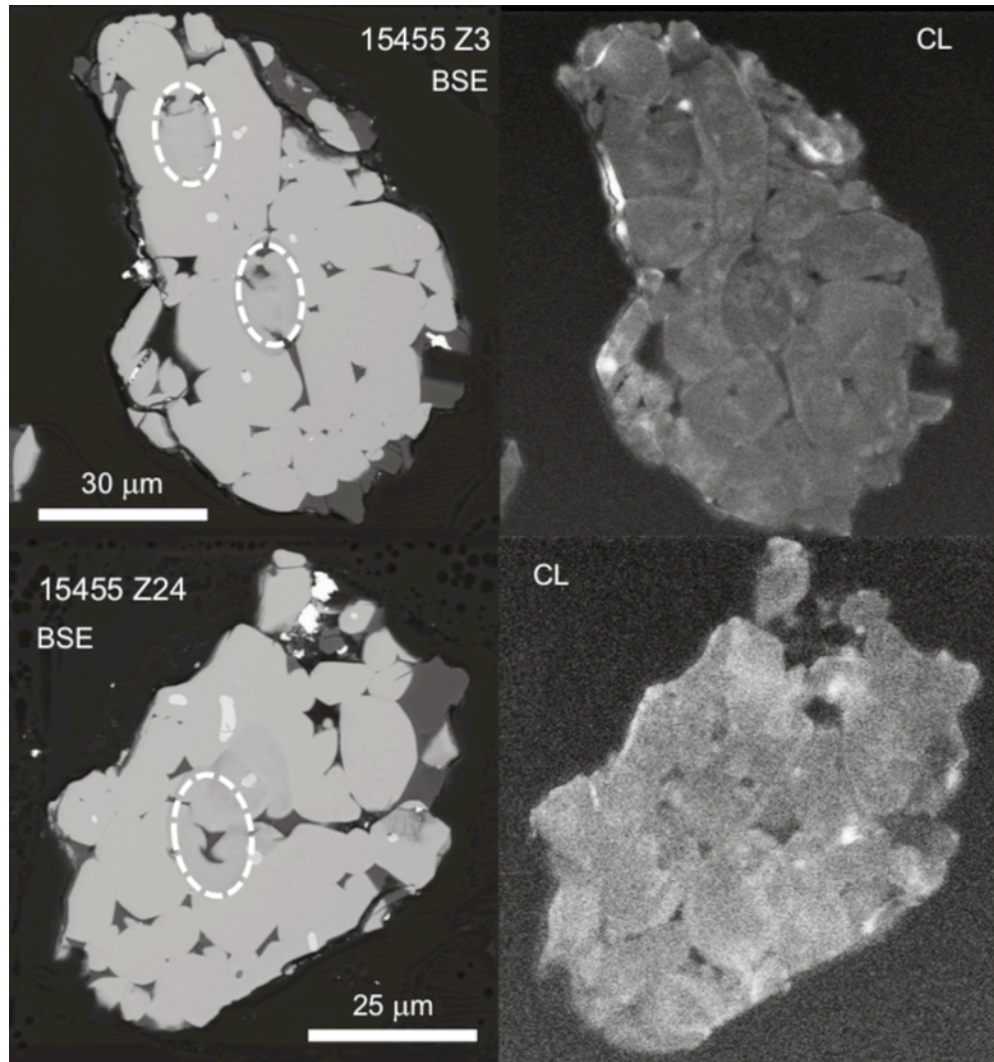
Interpreting lunar ages



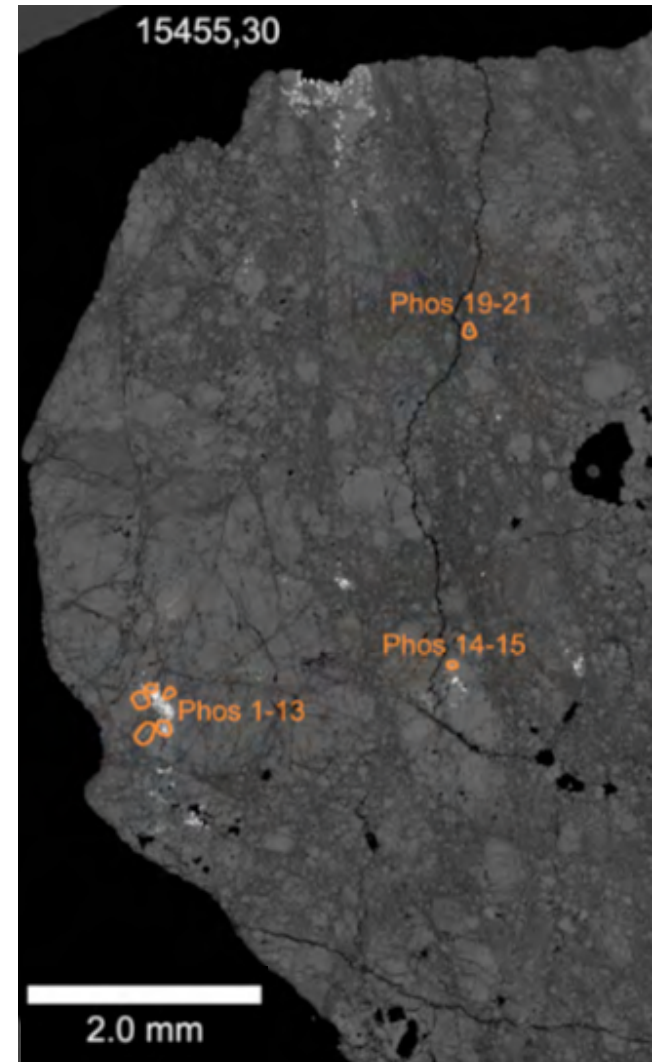
Lunar zircon record



Zircons and phosphates from 15455



Crow et al., 2017



Nemchin et al., 2020